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Abstract	<p>This chapter examines the human search, understanding, and interpretation of biosignatures. It deals with four epistemological issues in the search for signs of life in outer space: (1) conceptualization, how we form concepts of life in astrobiology, how we define and categorize things, and the relation between our concepts and our knowledge of the world; (2) analogy, how we see similarities between things, and with inductive, analogical reasoning go from what we know to what we do not know, from the only example of life here on Earth, to possible extraterrestrial life; (3) perception, how we interpret what our senses convey in our search for biosignatures, how the information we get from the surrounding world is processed in our minds; and (4) the semiotics of biosignatures, how we, as interpreters, establish connections between things, between the expression (the biosignature) and the content (the living organism) in various forms of semiosis, as icons, indices, and symbols of life. In all, it is about how we get access to the world, and how we interpret and understand it, for achieving a well-grounded knowledge about the living Universe.</p>
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Chapter 15 1
The History and Philosophy 2
of Biosignatures 3

David Dunér 4

... all this universe is perfused with signs, if it is not composed exclusively of signs. 5
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C. S. Peirce, "The Basis of Pragmatism in the Normative Sciences" (1906, 1998, 394). 8
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Abstract This chapter examines the human search, understanding, and interpretation of biosignatures. It deals with four epistemological issues in the search for signs of life in outer space: (1) conceptualization, how we form concepts of life in astrobiology, how we define and categorize things, and the relation between our concepts and our knowledge of the world; (2) analogy, how we see similarities between things, and with inductive, analogical reasoning go from what we know to what we do not know, from the only example of life here on Earth, to possible extraterrestrial life; (3) perception, how we interpret what our senses convey in our search for biosignatures, how the information we get from the surrounding world is processed in our minds; and (4) the semiotics of biosignatures, how we, as interpreters, establish connections between things, between the expression (the biosignature) and the content (the living organism) in various forms of semiosis, as icons, indices, and symbols of life. In all, it is about how we get access to the world, and how we interpret and understand it, for achieving a well-grounded knowledge about the living Universe. 11
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28 **15.1 Introduction**

29 In the quest for life in Universe, the most likely scenario is that we one day might
30 find signs of life, biosignatures, that indicate certain biochemical processes that
31 could have their origin in extraterrestrial biological activity. Probably, this is what
32 we can hope for, inasmuch as we will not in the foreseeable future find ways of
33 exploring, in situ, foreign worlds around other stars. With current technology no
34 manned mission will take us there. The hope of finding life on other planets or
35 moons in our own Solar System, though, has still not vanished. In some hidden
36 corner of our Solar System some microbes or unicellular organisms might hide, but
37 some more complex forms of life would be very unlikely to find. However, it
38 becomes more and more clear, that among the hundreds of billions of stars in our
39 galaxy there are perhaps many millions of Earth-like planets of which many might be
40 habitable and have the right conditions for harbouring life. Even though we will
41 never hold these actual life-forms in our hands, nor be able to construct an optical
42 telescope that will let us get a glance of the planetary surface, it is fully conceivable
43 that we in the near future will be able to refine our methods and observations in order
44 to find signs of life.

45 On one hand we might come across observable and verifiable phenomena that we
46 call “biosignatures,” and on the other, we infer the existence of certain unknown
47 instances of known biochemical processes that we call “life” that we suppose are the
48 causes of the former. In other words, we make connections between the expression
49 (the biosignature) and the content (the living organism). The ones who make this
50 connection are we human beings, with our inventive minds that are a result of a
51 particular bio-cultural coevolution of human cognition, of our species, here on Earth
52 (Dunér and Sonesson 2016). This is what the history and philosophy of
53 biosignatures is all about. We cannot, by no means, rule out ourselves in the
54 inference. The data obtained, even if processed by a computer, must be
55 comprehended and interpreted by the human mind to acquire significance. The
56 interpreters of the “biosignatures” are and will be we. No one will help us with the
57 interpretation. In that respect we are alone in the Universe and need to rely on our
58 own interpretive capabilities. For us to be successful in our search for life beyond our
59 Solar System depends on, first, obviously, that it really exists life to be discovered;
60 secondly, that we have the technology to discover it. But, that is not enough. The
61 discovery depends on—and that is the most critical thing—the capability of the
62 human brain, the organization and efficiency of that systematic search for knowledge
63 that we call science, which is a product of the socio-cultural history of our species
64 (Dunér 2016a). Our endeavour depends on human cognition and our ability to
65 understand and interpret what we observe in our surrounding world.

66 Here is not the place for discussing the actual scientific research about
67 biosignatures per se; it is dealt with elsewhere in this volume. Anyhow,
68 biosignatures concern various things, and refer to chemical substances (elements,

molecules, etc.), but also physical features (structures, shapes, morphology, etc.), 69
and physical phenomena (electromagnetic radiation, light, temperature, etc.). They 70
can vary in scale from atomic to planetary magnitude, or perhaps even larger. They 71
can be searched for both by in situ investigations and through remote indirect 72
sensing, on our nearest planets and moons as well as in other Solar Systems. 73
These signatures are meant to be evidence for either living life or dead life, present 74
or past life, distinctive from an abiogenic background. What I aim at here is to 75
examine the human endeavour to make sense of the world around us, how we think 76
about biosignatures, rather than explaining what they are. In the following, I will 77
delve into the epistemic questions that the search for biosignatures provokes. The 78
epistemology of astrobiology is a less explored philosophical territory concerning 79
the limits of astrobiological knowledge, i.e., what is known, what is knowable in 80
practice or in principle, and what is knowably unknown (Dunér 2013a; cf. Persson 81
2013). The epistemological problems of astrobiology are somewhat similar to those 82
of other branches of science, but with the exception that the limits of our 83
astrobiological knowledge seem to be much more uncertain. In this chapter, I will 84
discuss four epistemological issues in the search for biosignatures: (1) conceptualiza- 85
tion, how we form concepts of life in science, how we define and categorize 86
things, and the relation between our concepts and our knowledge of the world; 87
(2) analogy, how we see similarities between things, and with inductive, analogical 88
reasoning go from what we know to what we do not know; (3) perception, how we 89
interpret what our senses convey, how the information we get from the surrounding 90
world is processed in our minds; and (4) the semiotics of biosignatures, how we, as 91
interpreters, establish connections between things, between the expression and the 92
content in various forms of semiosis, as icons, indices, and symbols of life. In all, it is 93
about how we get access to the world, and how we interpret and understand it, for 94
achieving a well-grounded knowledge about the living Universe. 95

15.2 The Concept of Life 96

When searching for biosignatures, the first and foremost challenge is to determine if 97
the sign is of a biotic or abiotic nature. Are there any ways of distinguish life from 98
nonlife? This differentiation between biotic and abiotic leads us to the question “What 99
is life?” If we are searching for something that we call “life,” we should at least know 100
what kind of phenomenon we have in mind. This seems perhaps obvious, but 101
however, it is trickier than what we at first glance might expect. The definition of 102
life is one of the most debated and discussed philosophical questions in astrobiology. 103
No consensus has so far been established. In the course of the history of the philosophy 104
of biology amounts of definitions have been put forward, one more inventive and 105
clever than the other, but again, none seems to be exhaustive and indefectible (e.g., 106
Luisi 1998; Cleland and Chyba 2002, 2007; Ruiz-Mirazo et al. 2004; Oliver and Perry 107

108 2006; Robus et al. 2009; Bedau and Cleland 2010; Gayon 2010; Pross 2012; Losch
109 2017). When examining Earth-like life, we find that it displays a number of charac-
110 teristics inherited due to a common origin: it is carbon-based, uses a few specific
111 organic molecules, and further more it is something that we perceive as alive, have
112 some sort of energy consumption, metabolism, and that it to some extent grows, and
113 transforms. One can make up a shortlist of ecological requirements for life (McKay
114 2007), e.g., energy, carbon, liquid water, and other elements such as nitrogen,
115 phosphorus, and sulphur. Life is made of these and other chemical components,
116 which are related to the surrounding environment, but in different proportions (Conrad
117 2007). Campbell and Reece (2002) and Domagal-Goldman et al. (2016) have listed a
118 number of traits that are common to life on Earth: ordered structure, reproduction,
119 growth and development, energy utilization, response to the environment, homeostasis
120 (to maintain a steady internal environment regardless of the external environment), and
121 evolutionary adaptation. But which are the most constitutive attributes of life, repro-
122 duction, evolution, metabolism, deoxyribonucleic acid, entropy resistance... or what?
123 Definitions of life that have been put forward commonly combine especially metab-
124 olism, reproduction, and evolution as the most decisive attributes (Palyi et al. 2002).
125 The most popular definition, NASA's "working definition," defines life as "a self-
126 sustaining chemical system capable of Darwinian evolution" (Joyce 1994). As has
127 been noted (Domagal-Goldman et al. 2016), its strength is that it stresses on life as an
128 evolutionary process, rather than its chemical composition. It pinpoints life as an
129 evolutionary process in contrast to the individual sample of life that does not undergo
130 Darwinian evolution itself. However, one could question how such a definition would
131 be useful when searching and analysing biosignatures. Could the life that the
132 biosignatures refer to be tested and proven to be capable of Darwinian evolution?
133 Perhaps more helpful for a hunter of biosignatures is the first part of the definition. A
134 "self-sustaining chemical system" should in one way or another differ from its
135 surroundings. In all, a definition should tell us if the phenomenon we encounter is
136 life or not, but also be broad enough so we will not dismiss samples of "life" that are
137 not similar or identical to terrestrial life.

138 The difficulty in arriving at an acceptable definition has to do with, among other
139 things: (1) what we mean by "definition" and what such a definition should do for us
140 in our search for life or what it should explain (cf. Persson 2013); (2) how we
141 categorize things and how concepts are formed and used in our minds; and (3) that
142 we know only one living planet in Universe, our own planet, and we still do not know
143 how life emerged here on Earth. The aim of the following is not to fully scrutinize the
144 philosophical question of the definition of life, even less to arrive at a definite
145 definition. But before we go into the epistemic and semiotic questions involved in
146 the search for biosignatures, which are the main target of this chapter, we need to
147 delve into the very concept that is in the focus for our search, life, and how it is
148 connected to our understanding of definitions, categorization, and our own ignorance.

15.2.1 *The Definition of Definition*

149

Science, as well as in the case of astrobiology, concerns concepts. We need names and abstract concepts in order to be able to talk and reason about objects, structures, processes, etc., that we gather from our senses, through observations and experiments. When we are using these terms, they need to be reasonably well defined, some sort of consent needs to be established, so we can agree on what we talk about. We need definitions. One might think that these concepts are already out there, just for us to discover, but the input we get from the surrounding world has to actually be processed by our brains and depends on the cognitive abilities we possess. The scientific concepts we use are not just dependent on the particular characteristics of human cognition that is a product a biological evolution of the human brain, it is also a product of a specific cultural evolution here on Earth and many generations of natural philosophers and scientists in the history of human, terrestrial science.

Constructing concepts in order to be able to think and talk about the new phenomena encountered is a major task for astrobiological research. Astrobiologists use a wide range of concepts, besides biosignature, for example life, habitability, habitable zones (e.g., Kane and Gelino 2012), Earth analogues, exoplanets, and other concepts and terms inherited from already well-established scientific disciplines, which together form that multidisciplinary field of research we call astrobiology. The most debated and discussed concept in astrobiology is, as said, the concept of life. However, my point here is not how we actually define life, but rather what we mean with “definition” and what it should do for us. There are a number of ingredients life needs to have in order to be life, as mentioned above, but if life is a recipe, what are the essential ingredients and which are optional ones? So far, the debate has intuitively employed an Aristotelian conception of definition (Aristotle 1966, 2.3.90b30–31), in which a “definition” is a limited list of characteristics that are necessary and/or sufficient for something to be of the type of object it is, and from which all the characteristics of the object originate. Many definitions of life tend to be a list of necessary and/or sufficient properties that something needs to have in order to be called “life,” and further more, this list has the pretension to be complete. In our daily lives, however, we make relatively little use of Aristotelian definitions and depend much more on prototypes (Rosch 1975, 1978). Dogs, cats, and horses may seem to be more typical representatives for “life” than arsenic resistant microbes. In astrobiology, the prototype for life is terrestrial life, a self-sustaining chemical system capable of Darwinian evolution of that sort we find here on Earth. Another option is to comprehend concepts in Wittgensteinian terms, that there is a family resemblance among the essential features of the concept, a series of overlapping similarities, but no one common to all (Wittgenstein 1953; cf. Rosch 1987; Pennock 2012; Persson 2013), and in our case, no prototype, no typical representative of “life.”

189 The search for life in the universe has to a great extent highlighted how strongly
190 the evolution of life and environment are intertwined (Golding and Glikson 2011;
191 Schulze-Makuch et al. 2015; Cabrol 2016), that the coevolution of life and environ-
192 ment determines the uniqueness of an extraterrestrial life form (Watson 1999; Irwin
193 and Schulze-Makuch 2001; Kooijman 2004; Dietrich et al. 2006). A definition of life
194 should not only encompass all the life we know, it should also be anticipatory and be
195 prepared for putting future phenomena in one of the categories “life” and “nonlife.”
196 Future discoveries in astrobiology will most likely challenge our categorizations and
197 definitions, that is to say, our preconception of what the world is and not is. So, we
198 should be prepared to re-categorize and redefine our concepts. Future exobiological
199 systematics and taxonomy will face problems for how we identify, describe, and
200 categorize what we encounter. And in that respect, the taxonomy of future extrater-
201 restrial life will be a product of the human mind.

202 15.2.2 *The Categorization of Our Categories*

203 A definition of life aims to delineate life and nonlife, and be able to deal with the
204 intermediate evolutionary stages between nonlife and life. Our understanding of the
205 origin of life as a bio-chemical evolutionary process supposes that there are inter-
206 mediate stages from “dead” chemical elements to increasingly more complex forms
207 of life. But where can we draw the line and say to each other that after this point the
208 existing chemical congregations must be called life? Other entities appear as con-
209 ceptual intermediaries, such as viruses, that possess many characteristics similar to
210 organisms; they evolve, have an ordered structure, and go through maturation. But
211 for most biologists, they could not be regarded as life in a proper sense, due to their
212 lack of homeostasis, energy utilization, response to the environment, and perhaps
213 most critically, they cannot reproduce independently of the metabolism of their
214 infected hosts. Other biological entities, such as transposons, are also close to the
215 blurred borderline between life and nonlife. The evolution of life seems to suggest
216 that there is a continuum between life and nonlife, or grades of life, a series of
217 increasing possession of “liveability.” Could something be less or more alive? My
218 tentative answer is yes.

219 It is our human way of understanding the world, by using categories, that makes
220 problems for us when we are dealing with phenomena far beyond the well-known of
221 our daily lives. As said, the problem of defining life has to do with how we
222 categorize things (Lakoff 1990; Taylor 2003). All living creatures seem to categorize
223 the environment in terms of edible *versus* inedible, benign *versus* harmful, and so
224 forth. Categorization becomes more complex in human cognition. The human mind
225 tends to categorize, to see hierarchies and similarities between things, such as species
226 and genera in taxonomy (Berlin 1992; Dunér 2013b). When encountering new
227 unfamiliar things, the old categories, systems, and beliefs are challenged and

sometimes will fall short. Our thinking, science, and belief systems will have to be revised, which will lead to adjustments, adaptations, and compromises. But anyhow, we need these categories and invent new ones in order to handle the world around us; otherwise we would fall into a chaotic abyss of unsorted impressions.

15.2.3 Empirical Ignorance

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Rather than defining “life,” one might aim for understanding “life.” In astrobiology there is an endeavour to achieve an objective concept of “life” that can guide scientific research. The question is, if such objectivity is reachable. The way towards an objective, scientific concept of life is not, by no means, straightforward. A major question concerns how to distinguish those general characteristics of “life” in Universe from those that are specific to our own life here on Earth (Gayon 2010). Without additional examples of life, as Cleland and Chyba (2002) argues, it would be impossible to know if our concept of life refers to a universal, objective natural phenomenon or is just a subjective category. Some characteristics of life might be universal in its proper sense, that there are certain necessary features that life needs to have in order to be alive. The lack of a second instance of life makes it hard to distinguish those characteristics that are universal of life from those that are inherited due to a common origin. This empirical ignorance prevents us from arriving at a definite definition. If future search for life will be successful, it will fundamentally change or concept of life.

Searching for biosignatures is to a large extent a hypothetico-deductive endeavour. We formulate a hypothesis, that certain signatures indicate biological activity, and then we search for these “biosignatures,” make observations, and deduce that they indicate that life exists on that particular planet. However, only when all other hypotheses have been disproved, and only the habitation-hypothesis remains, then we can consider it to be established scientific knowledge. But in most cases biosignatures are not unambiguous. Both biological and abiotic processes can produce them. The question is how to distinguish true biogenic signatures from abiotic mimics. On one hand we need to avoid “false positives” that mimic life, and on the other hand we need to avoid “false negatives,” that real signatures of life are overlooked (Tarter et al. 2007; Horneck et al. 2016). Researchers have faced this ambiguity of biosignatures in the search for the earliest traces of life on Earth (Pilcher 2003; Golding and Glikson 2011; Westall and Cavalazzi 2011) as well as in the analysis of Martian rocks (Westall et al. 2015). A famous example of ambiguous experimental results is the controversy of the Viking mission results (Klein et al. 1976; Levin and Straat 1976, 1979, 2016). Viking 1 landed in Chryse Planitia and Viking 2 in Utopia Planitia. Both had instruments, such as gas chromatograph-mass spectrometers, for performing experiments in situ to detect traces of life, to search for organic molecules that Earth-based organisms have,

267 and gases that are consumed or produced in the metabolism of terrestrial organisms.
268 Despite a well-equipped mission, the results became ambiguous. Partly, this ambi-
269 guity of biosignatures lies within the human mind, in our definitions, categories, and
270 lack of empirical knowledge. Next, I will go on to another cognitive peculiarity of
271 human reasoning that plays tricks on us in our endeavours: analogical arguments.

272 15.3 The Analogy of the Earth-Twins

273 Above, I have put forward some of the reasons why the scientific community has
274 failed, and probably will fail, to come up with a consistent definition of life as a
275 limited list of necessary or sufficient conditions. Such a definition might, though, be
276 of heuristic value, i.e. as a practical, useful aid by analogy from the known. To begin
277 with, we assume that there is only one physics, one chemistry in the Universe, and
278 accordingly all phenomena will follow the same natural laws. We know just one
279 inhabited planet, and everything that live on that planet is related, have one common
280 origin. In astrobiology it is assumed that this particular planet, except that it is a
281 living planet, has no exceptional characteristics. There are billions of stars of the
282 same type as our Sun in the Universe, with presumably millions, billions of Earth-
283 like planets that possess similar physical characteristics as our planet. Earth is a
284 rather mediocre place in the Universe. This so-called mediocrity principle states that
285 there is nothing remarkable with Earth. If this is true, it opens up for the quest for
286 Earth-twins.

287 A common form of argumentation in astrobiology is analogical reasoning from
288 what we know to what we do not know (Dunér 2013c). An analogical argument
289 could be explained as a search for similarities, i.e., a way of selecting features in the
290 source domain that are to be mapped onto the target domain, and of transferring
291 relevant properties from the source to the target. If x has the properties P_1, P_2, P_3, P_4
292 $\dots P_n$, and there is a y that has P_1, P_2, P_3 , we may conclude that it also has P_4 . If we
293 know there is an x that has these qualities, and we discover a y that also has some of
294 these qualities, then we conclude that all y also have the quality that we are seeking,
295 P_4 , or formulated in first-order logic: $\exists x(P_1x \wedge P_2x \wedge P_3x \wedge P_4x \dots P_nx) \wedge \exists y(P_1y \wedge$
296 $P_2y \wedge P_3y) \Rightarrow \forall y(P_4y)$. The challenge is then to select the correct and relevant salient
297 features from an infinite number of possible ones in the source domain, features that
298 then will be transferred to and mapped onto the target domain. In some sense,
299 astrobiology as a whole is one single, great analogy. Starting from the one particular
300 type of life we happen to know something about, namely life on Earth, we proceed to
301 search for life on other planets. We predominantly look for life as we know it:
302 something needing liquid water, being based mainly on carbon, inhabiting a planet
303 of a certain magnitude, gravitation, atmosphere, chemistry, temperature, etc., that
304 revolves within the habitable zone, at a certain distance, period, eccentricity, incli-
305 nation, etc., of a solar-type star (a G2 main sequence star of 4.5 Gyr of age), which in
306 turn has to be of a certain size and temperature, and so on. There are a number of
307 problems with such an analogical argument: first of all, do we know all the necessary

and sufficient conditions for life? Another problem is that we restrict ourselves to one particular sort of life we happen to know, life as we know it, and might overlook other forms of “weird life,” life as we do not know (Davies et al. 2009). The third problem lies in the unpredictability of evolution, that there are some stochastic events involved in the evolutionary process. The question is if all ingredients of life are in place, will that inevitably, by necessity lead to the emergence of life? Is life a natural manifestation of matter? (cf. de Duve 1995).

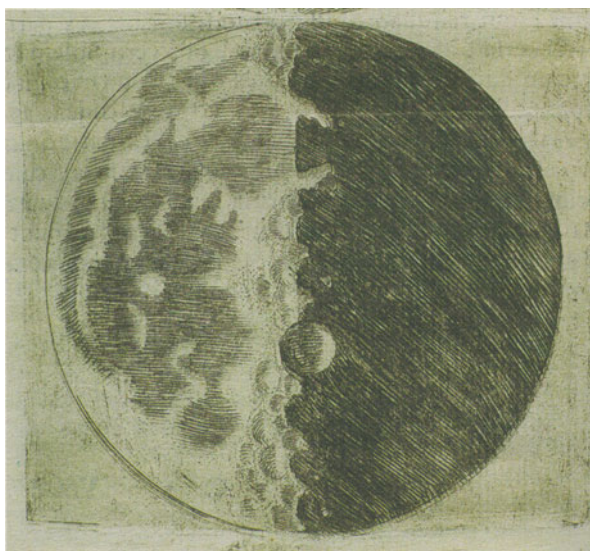
Logically speaking, analogical arguments are invalid. The historian of science Michael Crowe pointed out that one of the most pervasive logical and/or methodological fallacies in the plurality debate is mistaking necessary conditions for sufficient conditions (Crowe 1986). If liquid water is a necessary condition for life as we know it, it is just one among numerous other necessary conditions, which all need to be present if a planet is habitable. Just an evidence of an atmosphere, even if it contains the right chemical components including water vapour, is not enough to prove that life exists on its surface. All too often, Crowe states, finding some few necessary conditions among a larger set of necessary conditions, has been taken as a proof of life. If all necessary conditions need to be in place, the question arises if we know the complete set of conditions, or are there unknown conditions still to be discovered? The conclusion is that, not until we know all necessary and sufficient conditions for life, we will only be able to infer that the planet is habitable, but not that it is actually inhabited.

The fallacious use of analogical arguments were also discussed by the philosopher and logician Charles Sanders Peirce, stating that “There is no greater nor more frequent mistake in practical logic than to suppose that things that resemble one another strongly in some respects are any more likely for that to be alike in others” (Peirce 1957, 134; Crowe 1986). However, Peirce admitted, analogical argumentation could be accepted as a method of discovery, but not a method of proof. In providing us with some point of departure, analogical arguments still might hold some heuristic advantage in the search for life. Astrobiology as a great analogy is an inductive argument that cannot logically prove anything, just propose that life in Universe is a theoretical probability, but not an inductive probability. Just one sample of life, our own here on Earth, is a rather poor, to say the least, start of an empirical-inductive argument. Nor are large numbers, for example of stars, exoplanets, Earth-twins, any conclusive ground for an argument, just a theoretical probability.

What we are actually looking for is something that reminds us of ourselves, something similar to us, to earthly life, whether microbes or more complex life. In fact, to stretch it a bit, we are searching for “ourselves.” Though, life might be very different from what we imagine. The history of science is actually a history of surprises. The world we are living in turned out to be very different from what we first thought: richer, more complex, more peculiar and more astonishing than what we could dream of. This will also be true for astrobiology. Future discoveries in astrobiology will surprise us completely.

351 **15.3.1 The Mountains of the Moon**

352 Analogical reasoning is very common in the history of astrobiology and in the search
353 for biosignatures. Many works have discussed the history of the extraterrestrial life
354 debate and the emergence of the science of astrobiology (see e.g. Dick 1982, 1996;
355 Guthke 1983; Crowe 1986, 2008; Sullivan and Carney 2007; Dunér 2012, 2016b;
356 Briot 2013; Crowe and Dowd 2013; Dunér et al. 2016), but the concept of
357 biosignature has not got particular attention. In the extraterrestrial life debate of
358 the seventeenth and eighteenth centuries our closest celestial body, the Moon, was
359 the prime candidate for life on other worlds. A number of scientists and scholars also
360 speculated about life on Venus, Mars and on other planets, both within our Solar
361 System and beyond its frontiers.



November 30, 1609 Galileo trained his telescope on the Moon and, like the Earth, found it to be rugged and uneven, perhaps even having similar mountains and oceans. The first images of the Moon seen through a telescope were published in his book *Sidereus nuncius* (1610)

362 In the autumn of 1609, Galileo Galilei made the first closer observations of
363 extraterrestrial bodies through a telescope. In the *Sidereus nuncius* from 1610, he
364 shows, based on his telescopic observations and analogical reasoning, that the Moon
365 has mountains and therefore has the same solid, opaque and rugged nature as the
366 Earth. The irregular border between its dark and illuminated parts is incompatible
367 with the idea that it is a perfect spherical solid. Galileo wrote: “Anyone will then
368 understand with the certainty of the senses that the Moon is by no means endowed

with a smooth and polished surface, but is rough and uneven and, just as the face of the Earth itself, crowded everywhere with vast prominences, deep chasms and convolutions” (Galileo 1610; Spranzi 2004: 459). The Moon had a smooth appearance, though, in its contour, which he explained was because it might have an atmosphere. Galileo never stretched his analogical reasoning so far as he could by clearly claiming the existence of life on other planets. However, he did not consider it impossible that there were inhabitants on these spheres. And further more, we could not take it for granted that life elsewhere in the Universe must resemble our own. In 1612 Galileo wrote: “I agree with Apelles [the astronomer Christoph Scheiner] in regarding as false and damnable the view of those who put inhabitants on Jupiter, Venus, Saturn and the Moon, meaning by inhabitants, animals like ours, and men in particular” (Galileo 1957: 137; Dick 1982). Later in the *Dialogo . . . sopra i due massimi sistemi del mondo* (1632), he stated that there is no water, no humidity, no seas on the Moon, and therefore no life (Galileo 1632, 1953; Spranzi 2004).

In his *Cosmotheoros*, the Dutch scientist Christiaan Huygens (1698) expressed the view that it was highly probable that there was life on other planets. He noted that liquid water is necessary for life, and he saw darker and lighter spots on the surfaces of Mars and Jupiter that he interpreted as water and ice. Beyond our Solar System there are stars similar to our sun, and he asked why these stars could not have their own planets with their own moons. As for Venus, he empirically stated that a thick atmosphere surrounds it. He could not clearly detect any patches on the surface that might be signs of seas and mountains. Perhaps, he said, there are no seas on Venus, or, as he believed more probable, the air and clouds around Venus reflect nearly all the light from the Sun.

The first more certain telescopic observations of Venus, after Galileo’s discovery of its phases, were made during the 1761 transit of Venus. Many observers reported certain phenomena during the transit that they believed to have been caused by an atmosphere surrounding Venus (Proctor 1874; Woolf 1959; Maor 2000; Sellers 2001; Sheehan and Westfall 2004; Aspaas 2012; Wulf 2012; Sterken and Aspaas 2013). Certain astrodynamical facts relating to Venus were well known to the astronomical community, for example, its orbit around the Sun, magnitude, and phases, etc. However, the question of the atmosphere and topography of Venus was still unresolved. New observations of Venus against the solar disc changed the situation. The analogy argument started from the general supposition that there were no actual differences between Earth and Venus. They were both planets that orbited the Sun, were of similar size, solid, and, as some astronomers claimed, Venus also possessed mountains and an atmosphere. If there is life on Earth, then one may ponder why it could not also exist on Venus. If we can estimate the axial rotation and detect an atmosphere and mountains on that planet, it might also be true that it harbours life. These questions were in fact those that were investigated during the seventeenth and eighteenth centuries, and they included the length of its period of rotation and

411 whether it had mountains, an atmosphere and life. The Russian polymath Mikhail
412 Lomonosov argued that his observations of the transit of Venus in 1761 supported
413 the idea of a Venusian atmosphere: “Based on these observations I conclude that the
414 planet Venus is surrounded by a distinguished air atmosphere similar (or even
415 possibly larger) than that is poured over our Earth” (Marov 2005: 214f). Because
416 an existing atmosphere had been proved, then it could be concluded that Venus is
417 also inhabited.

418 Later Johann Elert Bode (1801) at the Berlin Observatory accepted astronomer
419 Johann Hieronymus Schröter’s (1793, 1796) claims about the existence of moun-
420 tains and valleys on Venus (Crowe 1986). Bode applied an apparently analogical
421 reasoning. He concluded that if Venus had land and sea, mountains and valleys,
422 clearings and condensations occurred in its atmosphere, and it had a companion
423 moon, then it was entirely similar to our Earth and consequently also habitable. The
424 French populariser of astronomy Camille Flammarion (1862) considered in *La*
425 *pluralité des mondes habités* that it was absurd that the Sun was employed solely
426 to illuminate and heat our small world. This absurdity became even more striking
427 when Venus was found to be a planet of the same dimensions as the Earth, with
428 mountains and plains, seasons and years, and days and nights analogous to our own.
429 That analogy was expanded to the conclusion, that, since they are alike in their
430 physical characteristics, they must be alike also in their role in the Universe: “if
431 Venus were without population, then the Earth must be similarly lacking, and
432 reciprocally, if the Earth were populated, Venus must be populated also” (Flamma-
433 rion 1862; Crowe 2008: 417f). Flammarion demonstrates here a characteristic
434 analogical thinking, typical of the astrobiological search for an Earth analogue. In
435 a later work on popular astronomy he says of the inhabitants of Venus: “this world
436 differs little from ours in volume, in weight, in density, and in the duration of its days
437 and nights. [...] It should, then, be inhabited by vegetable, human, and animal races
438 but little different from those of our planet” (Flammarion 1880, 1907: 371; Sheehan
439 and Westfall 2004).

440 Such analogical arguments could be summarized as a search for similarities in an
441 inductive manner, in order to pinpoint as many as possible, especially those of a
442 significant nature, i.e., those critical features that were believed to indicate a habit-
443 able environment. It was known that both Earth and Venus were planets of a similar
444 size, both orbited the Sun, and were exposed to its light and heat, and that both
445 globes were opaque and had a solid ground. These similarities could be extended, as
446 some astronomers maintained, to both of them having nearly exactly the same period
447 of axial rotation, as well as a companion moon, an atmosphere, mountains and seas.
448 If Earth and Venus seemed to be perfect twins, then there must be life on Venus too.

449 The last decades, analogy reasoning lies behind a number of experiments on
450 terrestrial life as preparation for in situ investigation on extraterrestrial bodies.
451 Terrestrial biosignatures have been treated as analogical models for possible extra-
452 terrestrial biosignatures. If we turn to ourselves, are we detectable; is it possible to

discover life on Earth from outer space? Based on observations made with the Galileo probe, Sagan et al. (1993) proposed what biosignatures might be possible to observe in the reflective light of Earth. According to the spectroscopic results, water, oxygen, ozone, carbon dioxide, carbon monoxide, nitrous oxide and methane were detected in the atmosphere of the Earth. The Very Large Telescope (VLT) of the European Southern Observatory (ESO) in the Atacama Desert, Chile, was used by Sterzik et al. (2012) for studying the Earthshine, the light from Earth that reflects on the surface of the Moon, seen as a greyish light on the lunar surface which is not lightened by the sun. The colour and polarization of the Earthshine showed that Earth's atmosphere contains clouds, that its surface is partially covered by a sea, and that it has vegetation. The idea was that by studying how Earth look like from space, one could get a reference for future analyses of exoplanet atmospheres. Analogy reasoning, helpful or not, rests on another cognitive process. In order to make an analogy between observations, one needs to interpret the perceptual information.

15.4 Epistemic Perception

The philosopher in Bernard de Fontenelle's *Entretiens sur la pluralité des mondes* (1686) states that all philosophy (understood as the natural sciences) is based on two things: curiosity and poor eyesight. We want to know more than what we can see. In contemporary astrobiology life seems always be beyond the fields we know. What the senses convey have to be interpreted through means of specific cognitive processes, and the interpretation of what has been observed is based on a preconceived understanding, concepts, and prior knowledge. Observations are not separate from theory. In contrary to the tendency to place excessive trust in "objective" observations, many philosophers of science, such as N.R. Hanson, Thomas Kuhn, and Michael Polanyi, have emphasized that observations are theory-laden, that there are no sharp dichotomy between observation and theory in scientific research (Crowe 1986). We need theories to understand what we see, and our preconceptions and expectations lead us in one or another direction, sometimes the wrong course. Not seldom we see what we expect to find.

In the optical observations of distant worlds, preconceived understanding often shapes the interpretation of what the observers see. Through their senses, the observers receive impressions from outer space, and they collect and collate information using their sight. What their senses convey have to be interpreted by means of specific cognitive processes before becoming reality. As observers, we do not just passively receive images and input from the world around us. Instead, the brain actively searches out patterns in what is conveyed to it through the senses, and interprets them through a process that is determined by both biological and cultural factors. Perception is not a neutral, objective, and realistic recording of reality. This conceptual or epistemic perception implies an identification of what is seen, and takes place by applying our concepts to visual perceptions, that is, concepts affect what we see, and, should we lack any concept of a specific phenomenon, then it will

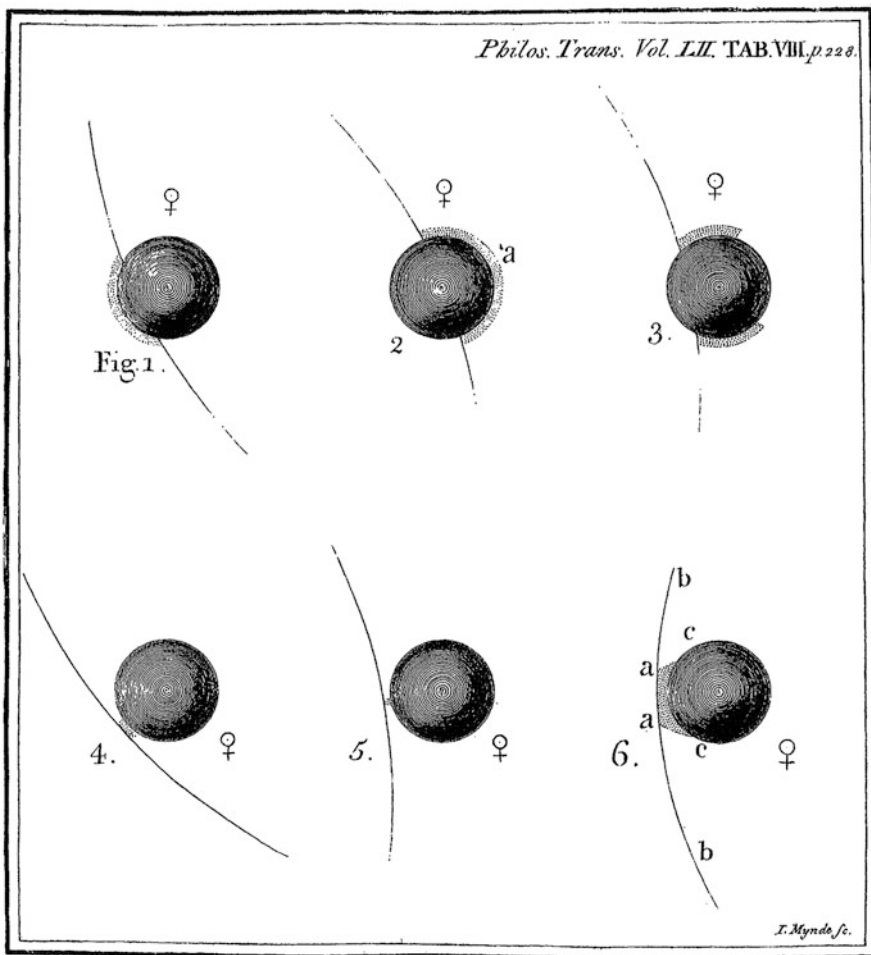
494 be difficult to distinguish it among all our impressions. The world distorts our
495 concepts, and the concepts distort our world.

496 **15.4.1 *The Atmosphere of Venus***

497 Astronomer William Herschel's lunar observations from 28 May 1776 highlight the
498 difficulty to see and arrive at a conclusive interpretation of the seen. He saw growing
499 substances: "My attention was chiefly directed to Mare humorum, and this I now
500 believe to be a forest, this word being also taken in its proper extended signification
501 as consisting of such large *growing substances*" (Crowe 1986, 63). This forest would
502 require, he said, trees at least 4–6 times the height of ours. In 1778 he even suspected
503 that lunar craters could be towns of the Lunarians. He saw circular buildings on the
504 Moon: "I am almost convinced that those numberless small Circuses we see on the
505 Moon are the works of the Lunarians and may be called their Towns" (Crowe 1986,
506 65). Those certain luminous spots that occasionally could be seen on the dark side of
507 the Moon, Herschel explained as volcanoes in eruption. Church minister and science
508 teacher Thomas Dick later wrote that a more pleasing idea would be that they were
509 "some occasional splendid illuminations, produced by the lunar inhabitants, during
510 their long nights" (Crowe 2008, 262f.).

511 Other striking examples of this epistemic perception are the maps of Venus that
512 delineated the surface of the planet, which showed mountains and other geological
513 features. Even a dim light, faint spots and lines, and a companion moon seemed to
514 appear when Venus was viewed through a telescope. The astronomers interpreted
515 their obscure observations in line with their prior knowledge and their ideas of the
516 nature of the world, and they often found what they sought. If they believed in the
517 existence of mountains and an atmosphere on Venus, then they duly found them. In
518 1645 the Neapolitan astronomer Francesco Fontana recorded "a dark patch in the
519 centre of the disk" of Venus, which can be said to be the first attempt to note
520 surface details there (Fontana 1646; Moore 1956; Cattermole and Moore 1997). In
521 1667, the astronomer Giovanni Domenico Cassini saw "various bright and dusky
522 patches" from which he deduced the first estimated period of rotation of 23 h and
523 21 min (Cassini 1667). Another Italian astronomer, Francesco Bianchini, drew the
524 first "map" in 1726 recording oceans and continents (Bianchini 1728; Sheehan and
525 Westfall 2004). On mist-free days, at twilight, he saw rounded patches similar to
526 lunar craters, and from their movements, he deduced that the period of rotation of
527 Venus was 24 days and 8 hours. There is no doubt that these records of the surface
528 features of Venus were purely optical. Beside the fact that the optical quality of the
529 telescopes was not always reliable, and that weather conditions could considerably
530 influence the quality of the observations, there is obviously also an epistemic

perception that changes the interpretations of the seen. The uncertain observations 531
 by Fontana, Cassini, Bianchini and others were interpreted in a particular way. If 532
 they believed in the existence of oceans and continents on Venus, they searched for 533
 them and found them, because their prior knowledge and beliefs directed their 534
 attention towards certain interpretations. The illusion or fallacy in their perception 535
 did not lie primarily and only in the flaws in their optical equipment, but in their 536
 imaginative minds, the cognitive apparatus that processed their sensory 537
 impressions. 538



The bright ring and the black drop seen from Uppsala during the Venus Transit of 1761. From Torbern Bergman's letter published in *Philosophical Transactions* (1762)

539 During the transit of Venus of 6 June 1761, two surprising phenomena were
540 observed: a bright ring around Venus and the “black drop” during the contacts. The
541 ring was re-observed on 3 June 1769, and its causes were still being debated even
542 then, but it was unanimously taken as proof of the existence of an atmosphere on
543 Venus. Concerning the black drop, the astronomer Daniel Melanderhjelm explained
544 it as caused by an atmosphere. The secretary of the Royal Swedish Academy of
545 Sciences, Pehr Wilhelm Wargentin, on the other hand, saw it as a mere diffraction
546 phenomenon (Dunér 2013c). Melanderhjelm’s observations were, as he said, just
547 “fallaciæ visus,” optical fallacies. The physicist Johan Carl Wilcke performed a
548 number of experiments during the summer of 1769 showing that the same phenom-
549 enon arises with a black body seen against a luminous body without any need to
550 assume an atmosphere. Wargentin and Wilcke, however, did not disbelieve in the
551 existence of an atmosphere, but the black drop could not provide proof of it.

552 Whether Venus has mountains and a surface similar to the Earth with valleys and
553 seas had been debated ever since the first Venus maps appeared in the seventeenth
554 century. Here, again, the conclusions were often a result of analogical reasoning and
555 epistemic perception. The great observational astronomers Herschel and Schröter
556 engaged in a heated argument as to the presence or absence of mountains on Venus.
557 However, they both agreed that Venus has an atmosphere. It was well known in the
558 era of Schröter and Herschel that Venus and the Earth, with regard to their size and
559 mass, were almost perfect twins. It then became also reasonable to assume that their
560 atmospheres were similar too, with regard to their extent and composition. In
561 February 1788 Schröter perceived the ordinarily uniform brightness of the disk as
562 being marbled by a filmy streak, and he concluded that what he was seeing was the
563 outmost part of a dense, cloudy atmosphere (Cattermole and Moore 1997). More-
564 over, the horns of the crescent were seen to extend beyond the semi-circle, which
565 could not be the case in the absence of an atmospheric mantle. On 28 December
566 1789 Schröter saw that the southern cusp of Venus was blunted, and that there was a
567 small luminous speck beyond it (Baum 1973). He saw the same thing again in 1790
568 and 1791 and concluded from these observations that it must be a very lofty
569 “enlightened mountain” that was catching rays of the Sun. Herschel re-observed
570 Venus in 1793, and he questioned Schröter’s findings. He agreed that Venus has an
571 atmosphere, but he never found those high mountains that Schröter mentioned
572 (Herschel 1912). In the *Philosophical Transactions* he states: “As to the mountains
573 in Venus, I may venture to say that no eye, which is not considerably better than
574 mine, or assisted by much better instruments, will ever get a sight of them” (Herschel
575 1793, 216; Moore 1956; Baum 1973).

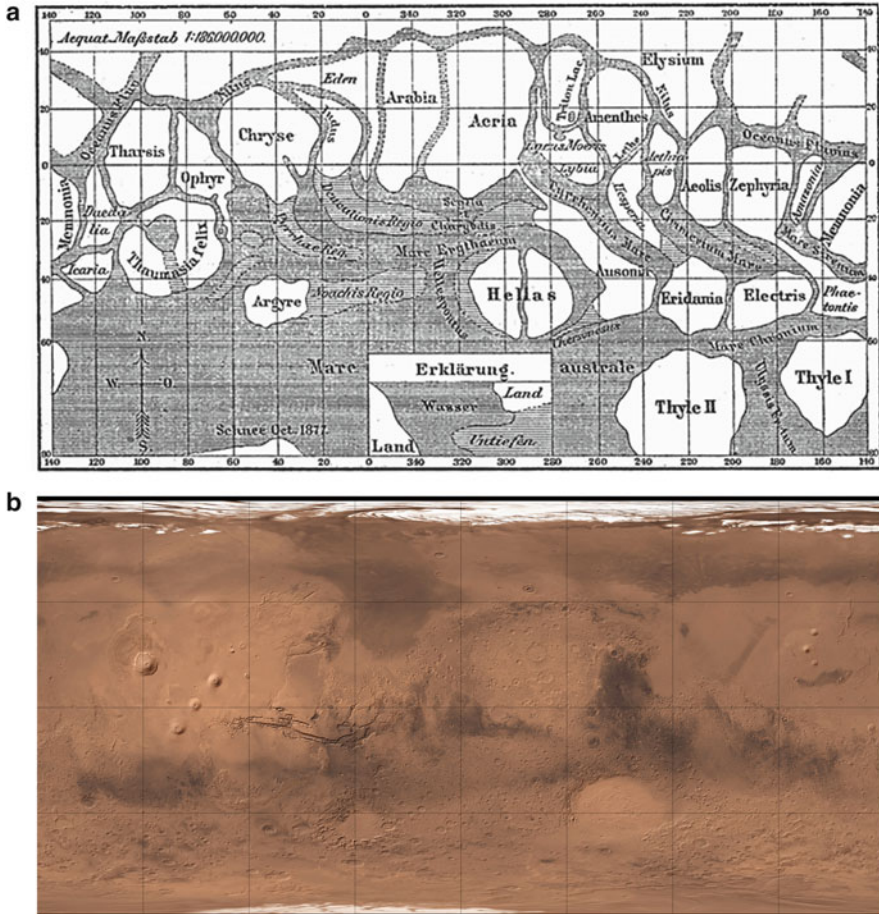
576 The observers of Venus saw things that needed explanations and interpretations.
577 They seemed to detect vague spots, streaks, lines, drops, a dim light, and a vague
578 companion. Such optical misinterpretations, or rather what could be explained as an
579 epistemic perception, were involved in the claims of a habitable Venus. The Ashen

light, the dim visibility of the non-sunlit side of Venus, when it is in the crescent stage, was first reported in 1643 by the Italian Jesuit and astronomer Giovanni Battista Riccioli. This phenomenon was also the object of an epistemic perception leading to certain interpretations of the seen. The German astronomer Franz von Paula Gruithuisen in Munich declared that the light had been seen in 1759 and again in 1806, an interval of seventy-six Venusian years, and he wrote: "The observed appearance is evidently the result of general festival illumination in honour of the ascension of a new emperor to the throne of the planet" (Cattermole and Moore 1997, 17). However, Gruithuisen later modified his explanation and instead of a Venusian coronation, he suggested that the light might be caused by the burning of large areas of jungle to create new farmland. In a paper from 1824, "Discovery of Many Distinct Traces of Lunar Inhabitants, Especially of One of Their Colossal Buildings", Gruithuisen even reports that he had observed cities, forts, a temple, and animal trails on the Moon (Crowe 1986; Sheehan and Dobbins 2001). Later in the nineteenth century, the invent of spectroscopic analysis brought hope that lines for oxygen and water vapour in the atmosphere of Venus could be detected, and some preliminary results seemed to support that. It was only in 1966 and 1967 that the Russian space probes Venera 3 and Venera 4 dived into the cloud shield, and a very hostile Venusian environment was discovered (Harvey 2007).

15.4.2 *The Canals of Mars*

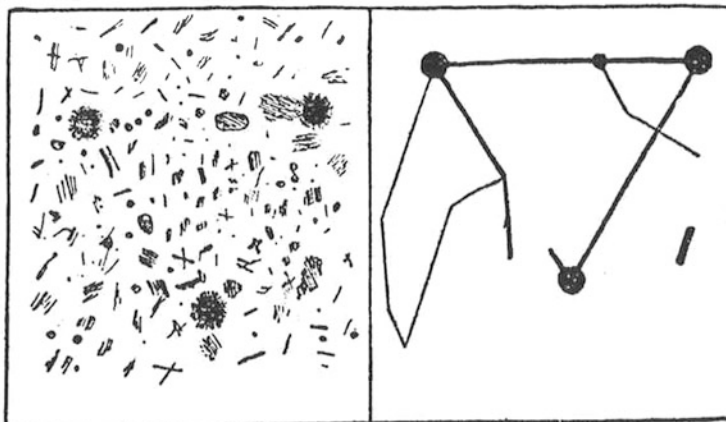
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Perhaps the most famous example in the history of astrobiology of such epistemic perception is the debate concerning the canals of Mars. In 1877 Giovanni Schiaparelli at Brera Observatory in Milan recorded in detail the Martian network of canals for the first time. In the beginning other astronomers had problem to confirm these observations, but during the next decades many succeeded. Schiaparelli's findings were confirmed by the American astronomer Percival Lowell who detected hundreds of Martian canals that he interpreted as an artificial irrigation system (Hoyt 1976; Strauss 2001). In his book *Mars* (1895) he claimed that he now had conclusive evidence of an intelligent civilization on the dying planet of Mars. A change of Martian coloration was also observed, as evidence for seasons and vegetation. A darkening of the surface spread when the spring came on.



Giovanni Schiaparelli's map from 1888 of the canals of Mars compared to a modern satellite map

611 But there were sceptics. Edward Walter Maunder of Greenwich Observatory
 612 questioned the canal observations as mere optical illusions. The eye tends to
 613 integrate details, he explained, that are below the limits of vision. Eugène Antoniadi
 614 used the high quality telescope at Meudon Observatory, and made a drawing of a
 615 specific region of Mars, which he compared with Schiaparelli's drawing from the
 616 same area. Where Schiaparelli's drawing showed canals, Antoniadi's saw just
 617 diffuse details, no canals. Schiaparelli's canals had been, according to Antoniadi,
 618 created by connecting dots on the surface or on the border between darker and lighter
 619 regions. In the American magazine *Popular Science* 1912, the reader could by
 620 himself experience this optical illusion with a simple, practical test consisting of
 621 random dots (Hennessey 1999). By then most astronomers had abandoned this
 622 imaginative idea.



Is Professor Lowell deceived? Hold this 20 feet away, and the marks on the left look like the lines on the right. From *Popular Science* (1912)

Mapping the planets involves great patience, to monitor the shimmering images 623
of the planets disturbed by the atmospheric conditions on Earth. The blurred images 624
seen in the telescope were faint sensory impressions that needed interpretations by 625
applying prior knowledge and conceptions to what the eyes perceived. The active 626
mind searched for regularities, order, and comprehensibility in the observations. The 627
story of human observations of the Moon, Venus, and Mars, tells us about the 628
imprecision of the human eye-brain-hand system under difficult conditions (Sheehan 629
1988; Sullivan and Carney 2007). When searching for biosignatures, we are looking 630
for regularities, order, and connections, based on previous knowledge, expectations, 631
and theories. Next, I will turn to the question how we connect the expression (the 632
sign of life) with its content (life). 633

15.5 Semiotics of Biosignatures

634

The semiotics of biosignatures is about the meaning-making processes of the human 635
mind and its ability to make meaningful connections between things. The problem of 636
biosignatures is very much a semiotic problem: how meaning can be discovered, 637
invented, deciphered, and interpreted. One might say that the science of astrobiology 638
“invents” connections between the signifier and the signified, expression and object, 639
“signs of life” and “life.” The first problem that arises in a situation of interpreting a 640
biosignature is realizing that it really is a sign at all. Some regularity and order, or 641
finding a repetition in the pattern is not enough. The sign should be recognized as 642
such by the interpreter, i.e., that it contains an expression that refers to a content, 643
leading to an interpretive process by the interpreter. In other words, the interpreter 644
needs to identify the physical phenomenon as containing semiotic meaning, 645

646 something that can be a sign of life, a biosignature, that has a particular meaning by
647 referring to its content “life.” In our everyday lives as well as in well-established
648 science, the expression of a phenomenon can easily be connected to its content. Seeing
649 the footprints of an animal, we can infer—based on previous knowledge—what kind of
650 animal it is, its weight and its direction. However, this semiotic confidence becomes
651 much more uncertain when we turn to biosignatures of unknown forms of life. We are
652 compelled to a number of assumptions that underlie the connection we try to establish
653 between the expression and the content. Our assumptions might be mistaken, which
654 leads to wrong interpretations, or even worse, we fail to make any assumptions at all, in
655 other words, the phenomenon encountered, “the signature,” would not be recognized as
656 a meaningful sign whatsoever. Even though we have good reasons to believe that the
657 connection we infer between the expression and the content, between the biosignature
658 and the living organism, is scientifically correct, we need to rule out other explanations
659 of the signrelation. The “biosignature” might not be a true biosignature at all, but
660 instead is caused by an unknown or known abiotic process.

661 Before we go further into the semiotics of biosignatures, it would be in place to
662 explain some fundamental semiotic concepts. As a study of signs, modern semiotics
663 has a long prehistory that could be traced back to the philosophy of Plato and
664 Aristotle, to John Locke and early modern medicine where semiotics was the
665 interpretation of bodily signs, to draw conclusion from symptoms, to arrive at
666 appropriate treatment of the patient. Particularly Peirce and Ferdinand de Saussure
667 in the turn of the last century formed semiotics (or in the latter’s case, semiology) of
668 today. A certain sub-field of research, biosemiotics, has focused on the production
669 and interpretation of signs in the biological realm (Wheeler 2006; Hoffmeyer and
670 Favareau 2008; Romanini and Fernández 2014). Cognitive semiotics, on the other
671 hand, studies the meaning-making structured by the use of different sign vehicles,
672 and the properties of meaningful interactions with the surrounding environment,
673 both with the physical and the social environment (Sonesson 2007, 2009; Zlatev
674 2015; Dunér and Sonesson 2016). Following attempt to uncover the meaning-
675 making strategies in the search for biosignatures takes its departure from cognitive
676 semiotics and related fields of research.

677 Semiotics is the study of meaning making. It concerns *signs*, which can be said to
678 be something that we interpret as having meaning. According to Peirce, one of the
679 main figures in semiotics, “A sign, or *representamen*, is something that stands to
680 somebody for something in some respect or capacity” (Peirce 1932, 135; Saint-
681 Gelais 2014). The sign, as *expression*, stands for something, its *object*. The sign does
682 not include its meaning, rather the meaning is attributed through elaboration of an
683 *interpreter*. So for something to be meaningful, an interpreter is needed, a human
684 being (or other meaning-making creatures) who endows the sign a meaning. A
685 physical phenomenon is meaningless so far as there is no one to recognize it as
686 meaningful. Phenomena that we call biosignatures become meaningful phenomena,
687 when we interpret them as containing a meaning by making a connection between
688 the expression and the object, in other words, between the “biosignature” and “the
689 living organism.” The signs are the way we make sense of the world, to approach it,
690 to get access to it and differentiate things from each other. Correspondingly, the

biosignatures we detect are one way among many other ways of making sense of the data we receive from outer space. And these biosignatures exist so far as we find them meaningful. In that perspective, the biosignatures are not solely “out there,” instead, they are to a great extent in our minds, in the interaction between our minds and the outer world. It is in our meaning-making practices the “biosignatures” become biosignatures. The *semiosis* is the sign process, how the signs operate in the production of meaning. In that sense, when the astrobiologist is interpreting biosignatures, he or she is involved in a meaning producing semiosis. This semiosis is triadic, it contains expression, object, and interpreter—which in our case respond to “biosignature,” “life,” and “astrobiologist.” Depending on how the interpreter makes or interprets the connection between the expression and the object, we have basically three types of signrelations, *icon*, *index*, and *symbol*. But, let us return to these signrelations in more detail later on. For now, it is enough to conclude that the astrobiologist searching for biosignatures is a sort of semiotician, an astrobio-semiotician, trying to establish connections between expressions and objects in the Universe. Semiotics of biosignatures concerns qualities and categories, as well as the search for rules and regularities within such a nomothetic science as astrobiology concerned with generalities. More generally, semiotics of biosignatures concerns the meaning-making processes of astrobiology.

Finding the connection between expression and content are actually mental, in the mind of the interpreter. The search for biosignatures is based on the human endeavour of connecting things with each other, and of selecting the right elements for the connection among a wider range of possible elements. We ask ourselves, what are the meaningful properties of the information we gather through spectrometers, radio or optical telescopes, etc.? Which signatures (phenomena) have meaning and which are just meaningless noises? Hence, we are looking for the meaningful signs among a chaotic “noise” of data, the “biosignatures” that clearly “says” that they are signs of “life.” The signifier is directly given, but the signified is only indirectly present, through the link with the signifier. “Life” (the signified) that we are searching for is just indirectly reachable for us, and we have to content ourselves with the only thing that is directly given, the biosignature (the signifier). As interpreters, we determine the relation between the signifier and the signified by picking out those elements we assume to be relevant. The challenge of the astrobiologist is to pick the right elements (properties) of the signifier. For example, when examining a Martian rock, we need to pick out those elements (shapes, molecules, etc.) that direct us to the signified, the living organism. And we need a reasonable explanation of the link between the signifier and the signified. Why, and how come, is this particular gas a result of a certain metabolism of a living organism? In what way is this shape a remnant of the morphology of a living organism? We need to know the physical processes that let us link the signifier with the signified.

In the following three sections, I will show that “biosignatures” is a very diverse category, not only in respect to its immense variety of expressions, but also in semiotic sense. It shows a great variation in signrelations that each has its particular epistemological problems. To begin with, in astrobiology the concept of “biosignature” is not completely unambiguous. The concept of “biomarker” is

736 sometimes used interchangeably with “biosignature,” but often restricted to refer to
737 “an organic molecule whose origin can be directly related to an organic component
738 of life” (Horneck et al. 2016, 227), or organic compounds characteristic of certain
739 organisms, for example hopanoids in cell membranes of cyanobacteria.
740 Biosignature, instead, refers to a broader range of signatures of life: morphological,
741 geochemical, and organic—a diverse and ambiguous category of signatures related
742 to either living organisms or fossilized remains. Biosignatures could be studied *in*
743 *situ*, being of physico-chemical, geological, morphological, and mineralogical
744 nature. Or they can be detected by remote methods, such as atmospheric spectroscopy,
745 chemical disequilibrium, isotope ratios, etc. (Hegde et al. 2015). Physical
746 microbial structures, stromatolites, mud mounds, atmospheric gases, etc., could be
747 biosignatures, varying in scale from prebiotic molecular features to entire planets,
748 referring to both life as we know and weird life, past and present life.

749 The following is a first attempt to bring some semiotic order in this chaotic variation
750 of signs. Based on Peirce three signrelations—icon, index, and symbol—one could at
751 least reveal some peculiarities of the semiosis of biosignatures. The meaning of the
752 relation between expression and content, that the interpreter experiences, is based on
753 either similarity (iconicity), proximity (indexicality), or habits, rules, or conventions
754 (symbolicity). An icon is when the expression shares some *similarity* with the object.
755 An index is when the expression has some *contiguity* with the object. And finally, a
756 symbol is when the connection between expression and object is just a mere convention.
757 Thus I put forward three kinds of biosignatures in semiotic sense: bioicons,
758 bioindices, and biosymbols. However, in many cases a sign could be a combination
759 of index, icon, and/or symbol, also as we will see in the case of biosignatures.

760 15.5.1 Icons of Life

761 Aristotle noticed similarities between certain seashell-shaped structures found in
762 rocks and those that washed ashore on the beach. Are there a connection between the
763 petrified seashells and the living ones? And how come? In the beginning of the
764 eleventh century, the Persian polymath Ibn-Sīnā (Avicenna) put forward the theory
765 of petrifying fluids as an explanation of the petrification. Through the ages, the
766 question was debated, if these figure stones that resembled living organisms actually
767 were fossilized seashells or just sports of nature, *lusus naturae*, if they grow in the
768 bedrock or were traces of once living animals exterminated by the flood. In 1665, the
769 Danish anatomist and geologist Nicolaus Steno found shark teeth in the Tuscan
770 mountains, suggesting that where it is now high mountains, it had once been a sea
771 (Cutler 2003). Other findings, however, seemed to have no counterparts in the living
772 species. By the end of the eighteenth century, the French paleontologist Georges
773 Cuvier began realizing that they actually were remnants of extinct species.

774 The idea of fossils was also combined with the idea of extraterrestrial life. If
775 meteorites were coming from outer space, as it was realized in early nineteenth
776 century, rather than being ejecta from volcanoes, these could be studied by chemical

and geological methods. These meteorites from other worlds might contain evidence 777
of extraterrestrial life, if not alive, in fossilized form. The analytical chemist Jöns 778
Jacob Berzelius (1834) discovered that meteorites contained organic materials 779
(hydrocarbons). He examined meteors from a meteor shower in November 1833, 780
made a chemical analysis of a carbonaceous chondrite that had fallen in 1806, in 781
Alais (Alès), France, but could not tell if it contained carboniferous of extraterrestrial 782
origin (Crowe 1986). In the 1870s it was consensus that some meteorites contained 783
organic materials, but no convincing evidence had been found that they contained 784
extraterrestrial life forms. With Berzelius and others the idea of panspermia, that 785
seed-bearing meteoric stones are moving around in outer space, became an increas- 786
ingly plausible area of research, ending up in the physical chemist Svante 787
Arrhenius's (1907) more elaborated panspermia hypothesis. And still, the hypothesis 788
has not been completely ruled out. Current research has shown that microbial life 789
could indeed travel between planets and survive in space (Horneck et al. 2010). 790

A theme in the history of palaeontology is the question of how to distinguish real 791
remnants of living organisms from structures that just mimic living forms, to 792
distinguish fossils from pseudofossils. These inorganic pseudofossils can be mis- 793
taken for fossils, for example branch-like structures like manganese dendrites in 794
limestone, kidney ore, moss agates resembling moss leaves and other patterns in 795
rock that arise through geological, not biological processes. This is still a challenge 796
in the quest for microfossils for tracing the early history of life on Earth or in order to 797
find fossilized life in Martian rocks. A famous example is the announcement in 1996, 798
that fossilized life had been discovered in the Martian meteorite ALH84001 (McKay 799
et al. 1996). Viewed under an electron microscope, certain tube-like structures in the 800
meteorite resembled fossilized bacteria. It was a premature claim, abiological pro- 801
cesses could in fact create these structures (Westall et al. 1998). This calls for new 802
samples when claiming evidence for fossilized life in rocks on Mars or beyond. A 803
second lot is needed to confirm or refute previous hypotheses based on the first batch 804
of samples, or to continue the search. 805

Biosignatures in the form fossils are distinctly another thing than remote sensing 806
of habitable atmospheres, not because how they are found, in situ, but in its 807
signrelation. Biosignatures that share a similarity with living organisms, for example 808
fossils, are in my terminology *bioicons*. In semiotics, an icon is a signrelation based 809
on similarity, where the expression shares some of the object's properties. The 810
similarity between properties is perceived on the background of other dissimilar 811
properties. The most obvious examples of bioicons are body fossils, the imprints of 812
the hard parts of animals and plants, where the imprints of skeletons or foliage let us, 813
based on morphologic similarity, establish a link between the fossilized structure and 814
the living thing. The very complexity of the expression (the fossil) directs us to the 815
conclusion, based on the supposition that such a complexity cannot be the result of 816
any known abiotic process. Microscopic fossils, microfossils, though, are more 817
challenging. All life as we know it share the characteristics of having internal 818
volumes isolated from the surrounding environment by a cell membrane. Based on 819
this shared morphology, one could search for cellular structures. Well-preserved 820
fossil cells can be identical in size, shape, and structure with living single-cell 821

822 organisms. Their structures, that show less complexity, make it however more
823 challenging to distinguish a biotic structure from an abiotic. On Earth, the Apex
824 chert from Western Australia, dated at ~3.5 Ga, has been claimed either to be
825 fossilized cells of filamentous bacteria or just a result of an abiotic process (Schopf
826 1993; Brasier et al. 2005). To be a true biosignature, it is not enough to notice a
827 similarity between the expression and the object. We also need an explanation for
828 how a living organism can become a fossil (a bioicon). If we find something that
829 reminds us of a living thing, a microbe, a cell or that like, we need also a theory that
830 links the living organism with the biosignature, establishing a physical correlation
831 between the bioicon and the thing it signifies. Plenty of performed experimental
832 fossilisation studies give us the right arguments to make this connection. Further
833 more, fossils are not enough if we want to get a complete understanding of the life it
834 refers to; they do not give us complete information of the biochemical nature of the
835 living organism.

836 Bioicons are not just of visual nature, a similarity based on morphology or
837 structure, they could exist in any sense modality. Based on chemical analyses, the
838 researcher sees similarities between the expression and the content, not because of
839 structural similarity, but because they share some chemical properties. In the case of
840 chemical biosignatures, some are bioicons in that sense that the discovered
841 biosignature has a chemical similarity or shared characteristics with the living
842 organism, for example complex biological macromolecules, like carbohydrates,
843 lipids, proteins, and nucleic acids (RNA and DNA). Most common biomolecules,
844 however, usually modify and degrade, and the products (also called “molecular
845 fossils”) of this chemical breakdown (the diagenesis) have instead an
846 indexical relation to the bioiconic biological macromolecules. For example the
847 2-methyl hopanes that are known to be the diagenetic products of
848 2-methylbacteriohopanepolyols are second order biosignatures, that is bioindices
849 of bioicons that refer to its content, life in the form of cyanobacteria.

850 15.5.2 *Indices of Life*

851 “It happen’d one day about noon, going towards my boat, I was exceedingly
852 surpris’d with the print of a man’s naked foot on the shore, which was very plain
853 to be seen in the sand: I stood like one thunder-struck, or as if I had seen an
854 apparition” (Defoe 1719, 122). When Robinson Crusoe, shipwrecked and washed
855 ashore on a seemingly uninhabited island, one day saw footprints in the sand, he
856 knew that there was human life there—Friday. He came to the conclusion, not just
857 because the footprint had a similarity (iconicity) with a human being, but because it
858 had a causal link with the life-form who made it, as an index of life. An index is a
859 sign caused by its object, it has an unintentional, causal link or contact with its
860 content. “Smoke”, for example, has this indexical relation to “fire,” which it refers to
861 and which we interpret as its cause. Indexicality is, in this respect, meaning by
862 proximity or contiguity. This contiguity does not necessarily have to be of real

physical causality, it could consist of the mere perceiving of two objects together in space. Indices could also be related to factorality, when seeing something as a part of something else (Sonesson 1994). The interpretation of indices requires empirical knowledge of the recurrent connection between the sign and what it refers to. The perceptual world consists of a profuse amount of potential indexicality, even though we do not yet recognize these indices as signs with meaning. But the human mind constantly searches for and infers causalities and meaning in things perceived.

Bioindices are thus biosignatures that have a connection to their objects (the living organisms) by contiguity. In other words, the connection between the expression and the content is not based on similarity, but on indexicality, and is in semiotic terms something distinctly different. Perhaps the clearest examples of bioindices are atmospheric, chemical biosignatures that refer to biological processes, such as the metabolism of living organisms. Homochirality and isotopic fractionation have been put forward as molecular evidence of metabolism. Biogenic minerals—deposits of calcium carbonate, calcium phosphate, iron oxides, manganese, and sulphur—could also be the products of microbial metabolic processes, and thus have this indexical relation, but are unfortunately very difficult to distinguish from minerals produced by mere abiotic processes. Fossils that record the behaviour or activity of an organism are another type of bioindices, in contrast to bioicons that has a similarity with the living thing. These artefacts of life, such as stromatolites formed by microbial mat communities, indicate a biotic origin. Other examples of bioindices that trace the behaviour of an organism are borings, burrows, footprints, etc. Again, the challenge here is to distinguish these bioindices from features that are a result of an abiotic process that mimic the biotic behaviour.

Remote sensing of planetary environments for habitability and biosignatures goes back to the nineteenth century. In his *Cours de philosophie positive* (1830–1842) the French positivist Auguste Comte said, concerning the celestial bodies, that “we will never by any means be able to study their chemical composition or their mineralogical structure” (Comte 1835, 2; Crowe 2008, 312). Some few decades later spectroscopy was developed. The turning point came with spectroscopic astronomy that gave a new powerful tool for searching extraterrestrial life. By analysing the spectra caused by the molecular absorption or emission at molecule-specific photon wavelengths, the spectroscopists could infer the chemical composition of the atmospheres of distant planets. The first spectroscopic observations aiming for detecting oxygen and water in the Martian atmosphere, were made by the astronomers William Huggins and Jules Janssen in the 1860s. By assuming water as a necessary condition for life, and by linking planetary environmental conditions (presence of water vapour in the atmosphere and liquid water on the surface) with the possibility for life to emerge and subsist, they got a clue. A detection of water vapour in the atmosphere of a planet would then be a crucial indication that it might be life on its surface. In 1867, Janssen claimed that he had discovered the presence of water vapour in the Martian atmosphere, but in fact it was probably terrestrial signatures, already refuted by the American astronomer William Wallace Campbell in 1894 (Raulin Cerceau 2013).

907 There are hopes that we in the future will be able to observe the absorption or
908 emission properties of atmospheres of small, rocky exoplanets (Seager 2014; Seager
909 and Bains 2015). A first step in the search for biosignatures of exoplanets would be
910 to study the temperature, size, mass, density, gravitation, and light conditions of the
911 exoplanet. Next, to search for indications of atmosphere, liquid water, clouds,
912 surface, plate tectonics, daily rotation, seasons, and weather. The third step would
913 be to look for bioindices. For sure, we will not be able observe any surfaces of
914 exoplanets with current technology, but we might soon be able to detect certain gases
915 that we connect with life by remote sensing, even though the interpretation of the
916 spectra involves a number of difficulties. In the future, the European Extremely
917 Large Telescope (E-ELT) will make it possible to perform spectroscopic analysis of
918 the faint light of an exoplanet, and might result in the first exoplanet atmospheric
919 biosignatures. Our hopes rest on the assumption that certain gases in the atmosphere
920 are produced by life (as we know from studies of our own terrestrial atmosphere),
921 such as oxygen, ozone, methane, and carbon dioxide. Oxygen enrichment in the
922 atmosphere could indicate the presence of oxygenic photosynthesis. Ozone, which is
923 produced photochemically from biologically produced oxygen, could be another
924 indication of biological activity. And methane could likewise be connected to the
925 metabolism of living organisms. However, these gases could also be produced by
926 abiological processes and exist without any biological activity. Some gases that are
927 products of life on Earth, such as CH_3Cl , CH_3SH , NO_2 , NH_3 , would not be detected
928 with current technology, due to low amounts, others, such as water and carbon
929 dioxide have significant abiotic sources, and are less suitable as conclusive signa-
930 tures of life. To conclude, the argument starts from the premises: (P_1) that life
931 produces certain gases as a by-product of metabolism; (P_2) some of these gases
932 will accumulate in the atmosphere; and (P_3) that these gases show a unique spec-
933 trum. From these premises—which we hold, to be true and to be sufficient for
934 detection—we conclude that life could, in theory, be detected through spectroscopy.
935 But if P_1 is false (there are metabolic processes that do not produce gases) or if P_2 is
936 false (these gases do not leak into the atmosphere) or we do not recognise the unique
937 spectrum (P_3), we will fail.

938 It might rather be the combination of gases and the quantity of them, that closer
939 reveals if there are life on the planet. Life leads to disequilibria, for example in
940 respect to atmospheric chemical composition, entropy, etc. Earth-like atmospheric
941 biosignatures disappear relatively quickly on a planet where life has ceased to exist.
942 If there is a certain amount of a biosignature gas, it needs to have a continuous
943 source. That gases in disequilibrium could be diagnostic for life was first suggested
944 by Joshua Lederberg and James Lovelock in 1965 (Lederberg 1965; Lovelock 1965;
945 Catling and Kasting 2007). And this atmospheric disequilibrium is detectable by
946 spectroscopy, as in the case of spectral analysis of Earthshine (Arnold et al. 2002;
947 Arnold 2008). The simultaneous presence of oxygen and methane indicates an
948 atmospheric disequilibrium that could be assumed as a spectral evidence of life.
949 The sustainable source of these gases, in this ratio, is life. The discovery of
950 significant amounts of methane in the atmosphere of Mars then implies that there
951 must exist a recent or current source, otherwise the methane would rather quickly

disappear. The source could be geological activity and water in the subsurface—or subsurface biology (Domagal-Goldman et al. 2016). As a recurrent theme in the search for life, the signs are ambiguous.

Another shipwrecked traveller in foreign territories, the ancient Greek philosopher Aristippus, was cast ashore on the Rhodian coast. But when he found geometrical figures in the sand, he became convinced that he had come to a land inhabited by civilized people (Vitruvius 1934). These Rhodian bioindices did not only indicate life, people, they indicated civilization. A certain class of bioindices could be categorized as *technoindices*, a second order index that indicate technology, which in its turn could indicate life. When analysing the spectra of exoplanets, one might find signs that do not have any known natural origin, such as industrial pollution, artificial molecules, for example pollutants like chlorofluorocarbons (CFCs), or other artificial traces of environmental disequilibrium that reverberates across the biosphere (Lin et al. 2014; Shostak 2015; Frank and Sullivan 2016), which we interpret as technoindices of advanced life forms that are able to artificially manipulate their environment. Monitoring the stars and planets in our galaxy we perchance come across signs of extraterrestrial civilizations revealed by their use of technology, for example radio emissions or other electromagnetic radiation leaking out from their planet voluntarily or involuntarily. Finding signs of technology does not necessarily lead to the conclusion that they originate from a civilization consisting of biological creatures, if one think of the highly hypothetic self-replicating “von Neumann machines” that replicate and disperse themselves without the dependence of biological creators.

The search for indices of life is a way of connecting phenomena around us, inferring that certain signs indicate a causal connection to their object and origin—life. This semiosis or meaning-making endeavour is however triadic, includes something more than expression and object, the biosignature and the living organism. As the astrophysicist Arthur Eddington (1920) touched upon: “We have found a strange footprint on the shores of the unknown. We have devised profound theories, one after another, to account for its origins. At last, we have succeeded in reconstructing the creature that made the footprint. And lo! It is our own” (Sullivan and Baross 2007, 6). It is ourselves, the interpreters that make this connection between expression and content. Searching for indices of life may reveal some knowledge about the living Universe wherein we live, but also an understanding of how we search for meaning in the seemingly chaotic world around us.

15.5.3 Symbols of Life

August 15, 1977, the Big Ear radio telescope in Ohio received a very strong narrowband radio signal that lasted for 72 seconds. While reviewing the record date, the astronomer Jerry R. Ehman was stunned and wrote the comment “Wow!” on the computer printout. This anomaly has not been confirmed nor repeated (Kraus 1979). The first problem one faces in such a situation is to determine if it is a natural

993 or an artificial signal, if one could rule out all known natural causes of the signal and
994 conclude that it is an artificial signal caused by an intelligent civilization with
995 advanced technology. This *technosignature* might indicate the existence of technol-
996 ogy, as such a technoindex. The next problem that arises is to determine if it is
997 something more than just an index of technology, but actually contains a message, a
998 content that is meant to be communicated, deciphered, and understood by the
999 receiver. Probably, we would not be satisfied with a mere conclusion that it is a
1000 technoindex, but that it contains symbolical information, that it is a *technosymbol*.

1001 Searching for extraterrestrial intelligence by means of radio astronomy has been an
1002 exciting challenge ever since the start of Project Ozma in 1960 (Sagan 1973; Weston
1003 1988; Tarter 2001; Drake 2011; Schuch 2011; Dunér 2015, 2017; Traphagan 2015;
1004 Vakoch and Dowd 2015; Cabrol 2016). The starting point of the argument is
1005 plausible. Electromagnetic leakage from Earth is detectable from outer space, and
1006 likewise, if an extraterrestrial intelligence is engaged in radio communication, we
1007 would be able, at least in theory, to detect its voluntary or involuntary broadcasts. The
1008 problem of interstellar communication lies not so much in the physical or technologi-
1009 cal constraints, even though they very much challenge our scientific and technologi-
1010 cal skills, but in the cognitive and semiotic problems that interstellar message
1011 decoding provoke (Vakoch 1998; Dunér 2011b, 2014; Sonesson 2013).

1012 Intelligence could be seen as an evolved mental gymnastics required to survive
1013 and reproduce within its specific environment. This includes the capability of
1014 representing activities and being able to make inner models of reality. By using
1015 symbols an intelligent creature could engage in abstract thinking detached from the
1016 environment, by which they can reason about things not existent; things that are not
1017 right in front of them, in a specific moment in time. Very effective tools for
1018 symbolizing thought are our communicational devices. According to the cognitive
1019 linguist John Taylor (2002), language can be understood as a set of resources that are
1020 available to the language user for the symbolization of thought, and for the commu-
1021 nication of these symbolizations.

1022 The problem with symbols is that they are conventional, or arbitrary, as the
1023 founder of semiology Ferdinand de Saussure (1916) called them. Icons and indices
1024 are signs that have some *non-arbitrary* similarity or contiguity with the signified, in
1025 contrast to the symbols' completely *arbitrary* relation. For example the word "life"
1026 has no causal link to what it stands for, nor does it resemble what it signifies. There
1027 are no intrinsic relationship between the expression and content whatsoever. It is the
1028 interpreters (the ones that construct the message and the ones that decode them,
1029 respectively) that joins them together and establish the connection between the
1030 expression and the content. And the matching between the transmitters' and the
1031 receivers' interpretation of the symbols is by no means self-evident. We may figure
1032 out the reference of the signal, but will probably have severe problems understanding
1033 extraterrestrial symbols. It is not impossible to imagine that the aliens would have
1034 certain knowledge about their environment that in its content is similar to our own
1035 knowledge of mathematics, physics, and chemistry. But their expression of it would
1036 most likely be very different from ours. It is the message's expression rather than its
1037 content that becomes the difficulty for the interpreter. In symbols, there is a gap

between the sign and meaning. Nothing in the physical appearance of the sign gives 1038
any clue to its object; they are instead linked by an arbitrary correlation. In fact, most 1039
attempts at interstellar message constructions violate this basic semiotic understand- 1040
ing of signs that distinguishes between expression and content. Symbols are 1041
detached representations and, as such, dependent on cultural and social interactions 1042
that create some specific regularities that have their origin in more or less stochastic 1043
habits, conventions, etc., of the species (Sonesson and Dunér 2016). Our communi- 1044
cation and symbolization have evolved through an evolutionary and cultural- 1045
historical process here on Earth, and are thereby constrained by our human bodies, 1046
terrestrial environment, and the socio-cultural characteristics of our species. And 1047
likewise, a potential information transfer containing a symbolic message from an 1048
alien civilization would be constrained by the bio-cultural coevolution of the extra- 1049
terrestrial intelligence that coded it. 1050

15.6 Conclusion 1051

So far, we have no conclusive evidence of the existence of extraterrestrial life. But 1052
could we ever be hundred percent sure that we are alone? One might object that the 1053
plurality of life hypothesis labours under a major problem: unfalsifiability. 1054
According to the philosopher of science Karl Popper (1963), theories that are 1055
unfalsifiable and rich in explanatory power are often known to be wrong. In our 1056
case, there is no method or way of proving that life does not exist and cannot exist 1057
elsewhere in Universe. Astrobiology has an attractive flexibility and opens up for a 1058
richness of various explanations for the failure to prove the existence or for future 1059
success in discovering life. If we do not find life on the surface of Mars, we continue 1060
searching under its surface. If we do not find life in our Solar System, we continue 1061
searching in other Solar Systems. If one exoplanet does not show any signs of life, 1062
we go on to the next, and so on. When will we give up and conclude that life most 1063
likely does not exist elsewhere? After hundred thoroughly studied exoplanets, after 1064
thousand, millions? We will never be able to search for life in every corner of the 1065
Universe, doing in situ investigations on every exoplanet in all galaxies in the entire 1066
Universe. We will never know empirically that life does not exist on other planets. 1067
We can just move on, refining our methods, observations, theories, etc., but never 1068
reach a final conclusion beyond uncertainty. It is just a question of probability. But to 1069
verify the hypothesis, it is just enough with one single sample. Finding life is not 1070
empirically falsifiable, but, in principle, verifiable. 1071

Anyhow, one day we might encounter signs of another living planet. We will then 1072
get some potential knowledge, among other things, about its chemical composition 1073
and environmental circumstances. But above all, the descriptions, interpretations, 1074
and conclusions concerning this new world will also tell something about ourselves 1075
and our place in the Universe, and how we interpret and understand that reality we 1076
experience. There are things we know. Even though life might not exist out there, it 1077
is we human beings with our brains, bodies, and cultures who are searching for 1078

1079 it. The history and philosophy of biosignatures is centred on humans, or more
 1080 specifically, on the scientific endeavour's dependence on the human mind and
 1081 human culture (Dunér 2011a, 2013a; Dunér et al. 2013). Astrobiologists have brains,
 1082 for sure; they are using cognitive tools that are a result of the bio-cultural coevolution
 1083 of human cognitive abilities. Certain cognitive processes are at work when astrobi-
 1084 ologists encounter unknown things, when interpreting potential signs of life, when
 1085 they gather and classify information, and make conclusions from the observational
 1086 data. This does not go on in subjective isolation. Astrobiologists live in a culture, in a
 1087 certain time in history, in a specific research environment, and collaborate with other
 1088 thinking beings. In this chapter, I have touched upon some epistemological issues in
 1089 the search for biosignatures, how we conceptualize things, make analogies, how we
 1090 perceive our surrounding world and endow it meaning. The quest for signs of life
 1091 rests on the cognitive and socio-cultural capabilities of that human species that
 1092 makes this lonely planet alive and thinking.

1093 References

- 1094 Aristotle (1966) Posterior analytics. In: Tredennick H, Forster ES (eds) *Topica*. Heinemann,
 1095 London
- 1096 Arnold L (2008) Earthshine observation of vegetation and implication for life detection on other
 1097 planets: a review of 2001–2006 works. *Space Sci Rev* 135:323–333
- 1098 Arnold L, Gillet S, Lardièrè O et al (2002) A test for the search for life on extrasolar planets: looking
 1099 for the terrestrial vegetation signature on the earthshine spectrum. *Astron Astrophys*
 1100 392:231–237
- 1101 Arrhenius S (1907) Panspermy: the transmission of life from star to star. *Sci Am* 196:196
- 1102 Aspaas PP (2012) Maximilianus Hell (1720–1792) and the eighteenth-century transits of Venus: a
 1103 study of Jesuit science and Central European contexts. Tromsø University, Tromsø
- 1104 Baum R (1973) *The planets: some myths & realities*. David & Charles, New York
- 1105 Bedau MA, Cleland CE (2010) *The nature of life: classical and contemporary perspectives from*
 1106 *philosophy and science*. Cambridge University Press, Cambridge
- 1107 Berlin B (1992) *Ethnobiological classification: principles of categorization of plants and animals in*
 1108 *traditional societies*. Princeton University Press, Princeton, NJ
- 1109 Berzelius JJ (1834) Ueber meteorsteine. *Ann Phys Chem* 33:113–148
- 1110 Bianchini F (1728) *Hesperii et Phosphori nova phaenomena, sive observationes circa planetam*
 1111 *Veneris*. Rome
- 1112 Bode JE (1801) *Allgemeine Betrachtungen über das Weltgebäude*. Berlin
- 1113 Brasier MD, Green OR, Lindsay JF et al (2005) Critical testing of Earth's oldest putative fossil
 1114 assemblage from the ~3.5 Ga Apex chert, Chinaman Creek, Western Australia. *Precambrian*
 1115 *Res* 140:55–102
- 1116 Briot D (2013) Elements for the history of a long quest: search for life in the Universe. *Int J*
 1117 *Astrobiol* 12:254–258
- 1118 Cabrol NA (2016) Alien mindscapes: a perspective on the search for extraterrestrial intelligence.
 1119 *Astrobiology* 16:1–16
- 1120 Campbell NA, Reece JB (2002) *Biology*, 6th edn. Benjamin Cummings, New York
- 1121 Cassini G (1667) Touchant la découuerte qu'il a faite du moueuement de la Planete de Venus à
 1122 l'entour de son axe. *Journal des Scavans*

- Catling D, Kasting JF (2007) Planetary atmospheres and life. In: Sullivan WT, Baross JA (eds) *Planets and life: the emerging science of astrobiology*. Cambridge University Press, Cambridge, pp 91–116 1123
- Cattermole P, Moore P (1997) *Atlas of Venus*. Cambridge University Press, Cambridge 1126
- Cleland CE, Chyba CF (2002) Defining “life”. *Orig Life Evol Biosph* 32:387–393 1127
- Cleland CE, Chyba CF (2007) Does ‘life’ have a definition? In: Sullivan WT, Baross JA (eds) *Planets and life: the emerging science of astrobiology*. Cambridge University Press, Cambridge, pp 119–131 1128
- Comte A (1835) *Cours de Philosophie Positive: Tome 2, Contenant la Philosophie Astronomique et la Philosophie de la Physique*. Bachelier, Paris 1131
- Conrad PG (2007) Instruments and strategies for detecting extraterrestrial life. In: Sullivan WT, Baross JA (eds) *Planets and life: the emerging science of astrobiology*. Cambridge University Press, Cambridge, pp 473–482 1133
- Crowe MJ (1986) The extraterrestrial life debate 1750–1900: the idea of a plurality of worlds from Kant to Lowell. Cambridge University Press, Cambridge 1136
- Crowe MJ (2008) *The extraterrestrial life debate, antiquity to 1915: a source book*. University of Notre Dame, Notre Dame, IN 1137
- Crowe MJ, Dowd MF (2013) The extraterrestrial life debate from antiquity to 1900. In: Vakoch DA (ed) *Astrobiology, history, and society: life beyond earth and the impact of discovery*. Springer, Berlin, pp 3–56 1138
- Cutler A (2003) *The seashell on the mountaintop: a story of science, sainthood and the humble genius who discovered a new history of the earth*. Dutton, New York 1142
- Davies PCW, Benner SA, Cleland CE et al (2009) Signatures of a shadow biosphere. *Astrobiology* 9:241–249 1143
- de Duve C (1995) *Vital dust: life as a cosmic imperative*. Basic Books, New York 1144
- de Fontenelle B (1686/1767) *Entretiens sur la pluralité des mondes*. Amsterdam: Mortier; trans. Conversations on the plurality of worlds. Thomas Caslon, London 1145
- de Saussure F (1916) *Cours de linguistique générale*. Payot, Paris 1146
- Defoe D (1719/2001) *Robinson Crusoe*. ed. Richetti J. Penguin, New York 1147
- Dick SJ (1982) *Plurality of worlds: the origins of the extraterrestrial life debate from Democritus to Kant*. Cambridge University Press, Cambridge 1148
- Dick SJ (1996) *The biological universe: the twentieth-century extraterrestrial life debate and the limits of science*. Cambridge University Press, Cambridge 1149
- Dietrich LEP, Michael M, Newman DK (2006) The co-evolution of life and earth. *Curr Biol* 16: pR395–pR400 1150
- Domagal-Goldman SD, Wright KE, Adamala K et al (2016) The astrobiology primer v2.0. *Astrobiology* 16:561–653 1151
- Drake F (2011) The Search for extra-terrestrial intelligence. *Philos Trans R Soc A Math Phys Eng Sci* 369:633–643 1152
- Dunér D (2011a) Astrocognition: prolegomena to a future cognitive history of exploration. In: Landfester U, Remuss N-L, Schrogl K-U, Worms J-C (eds) *Humans in outer space – interdisciplinary perspectives*. Springer, Wien, pp 117–140 1153
- Dunér D (2011b) Cognitive foundations of interstellar communication. In: Vakoch DA (ed) *Communication with extraterrestrial intelligence*. State University of New York Press, Albany, NY, pp 449–467 1154
- Dunér D (2012) The history and philosophy of astrobiology. *Astrobiology* 12:901–905 1155
- Dunér D (2013a) Extraterrestrial life and the human mind. In: Dunér D, Parthemore J, Persson E, Holmberg G (eds) *The history and philosophy of astrobiology: perspectives on extraterrestrial life and the human mind*. Cambridge Scholars Publishing, Newcastle-upon-Tyne, pp 7–31 1156
- Dunér D (2013b) The language of cosmos: the cosmopolitan endeavour of universal languages. In: Rydén G (ed) *Sweden in the eighteenth-century world: provincial cosmopolitans*. Ashgate, Farnham, pp 41–65 1157

- 1175 Dunér D (2013c) Venusians: the planet Venus in the 18th-century extraterrestrial life debate. In:
 1176 Sterken C, Aspaas PP (eds) Meeting Venus: a collection of papers presented at the Venus transit
 1177 conference in Tromsø 2012. *J Astronom Data* 19:145–167
- 1178 Dunér D (2014) Interstellar intersubjectivity: the significance of shared cognition for communica-
 1179 tion, empathy, and altruism in space. In: Vakoch DA (ed) Extraterrestrial altruism: evolution and
 1180 ethics in the cosmos. Springer, Dordrecht, pp 139–165
- 1181 Dunér D (2015) Length of time such civilizations release detectable signals into space, L, pre-1961.
 1182 In: Vakoch DA, Dowd MF (eds) The Drake equation: estimating the prevalence of extraterres-
 1183 trial life through the ages. Cambridge University Press, Cambridge, pp 241–269
- 1184 Dunér D (2016a) Science: the structure of scientific evolutions. In: Dunér D, Sonesson G (eds)
 1185 Human lifeworlds: the cognitive semiotics of cultural evolution. Peter Lang, Bern, pp 229–266
- 1186 Dunér D (2016b) Swedenborg and the plurality of worlds: astrotheology in the eighteenth century.
 1187 *Zygon J Relig Sci* 51:450–479
- 1188 Dunér D (2017) On the plausibility of intelligent life on other worlds: a cognitive-semiotic
 1189 assessment of $f_i \cdot f_c \cdot L$. *Environ Humanit* 9(2):433–453
- 1190 Dunér D, Sonesson G (2016) Human lifeworlds: the cognitive semiotics of cultural evolution. Peter
 1191 Lang, Bern
- 1192 Dunér D, Parthemore J, Persson E, Holmberg G (2013) The history and philosophy of astrobiology:
 1193 perspectives on extraterrestrial life and the human mind. Cambridge Scholars Publishing, New-
 1194 castle-upon-Tyne
- 1195 Dunér D, Malaterre C, Geppert W (2016) The history and philosophy of the origin of life. *Int J*
 1196 *Astrobiol* 15:1–2
- 1197 Eddington AS (1920) Space, time, and gravitation: an outline of the general relativity theory.
 1198 Cambridge University Press, Cambridge
- 1199 Flammarion C (1862) *La Pluralité des Mondes Habités*. Mallet-Bachelier, Paris
- 1200 Flammarion C (1880/1907) *Astronomie populaire: description générale du ciel*. Paris. In: Marpon
 1201 C, Flammarion E, trans. Gore JE, *Popular astronomy*. Appleton, New York
- 1202 Fontana F (1646) *Novæ observationes terrestriumq[ue] rerum observationes, et fortasse hactenus*
 1203 *non vulgatae*. Neapoli
- 1204 Frank A, Sullivan WT III (2016) A new empirical constraint on the prevalence of technological
 1205 species in the universe. *Astrobiology* 16:359–362
- 1206 Galileo G (1610/2009) *Siderevs nvnctivs magna, longeque admirabilia spectacula pandens, . . .*
 1207 *Venice*; trans. Shea WR, *Galileo's Sidereus nunciuss, or, a sidereal message*. Science History
 1208 Publications, Sagamore Beach, MA
- 1209 Galileo G (1632/1953) *Dialogo . . . sopra i due massimi sistemi del mondo . . . Firenze*; trans. de
 1210 Santillana G, *Dialogue on the great world systems*. Chicago University Press, Chicago, IL
- 1211 Galileo G (1957) *Discoveries and opinions of Galileo*, ed. & trans. Drake S. Doubleday, Garden
 1212 City, NY
- 1213 Gayon J (2010) Defining life: synthesis and conclusions. *Orig Life Evol Biosph* 40:231–244
- 1214 Golding SD, Glikson M (2011) Earliest life on Earth: habitats, environments and methods of
 1215 detection. Springer, Dordrecht
- 1216 Guthke KS (1983) *Der Mythos der Neuzeit: das Thema der Mehrheit der Welten in der Literatur-*
 1217 *und Geistesgeschichte von der kopernikanischen Wende bis zur Science Fiction*. Francke, Bern
- 1218 Harvey B (2007) *Russian planetary exploration: history, development, legacy, prospects*. Praxis,
 1219 Chichester
- 1220 Hegde S, Paulino-Lima IG, Kent R, Kaltenecker L, Rothschild L (2015) Surface biosignatures of
 1221 exo-earths: remote detection of extraterrestrial life. *Proc Natl Acad Sci USA* 112:3886–3891
- 1222 Hennessey R (1999) *Worlds without end: the historic search for extraterrestrial life*. Tempus, Stroud
- 1223 Herschel W (1793) *Observations on the planet Venus*. *Philos Trans* 83
- 1224 Herschel W (1793/1912) *Observations on the planet Venus*. In: Dreyer JLE, Herschel W (ed) *The*
 1225 *Scientific Papers of William Herschel*, London
- 1226 Hoffmeyer J, Favareau D (2008) *Biosemiotics: an examination into the signs of life and the life of*
 1227 *signs*. University of Scranton Press, Scranton

Horneck G, Klaus DM, Mancinelli RL (2010) Space microbiology. *Microbiol Mol Biol Rev* 74:121–156 1228
1229

Horneck G et al (2016) AstRoMap European astrobiology roadmap. *Astrobiology* 16:201–243 1230

Hoyt WG (1976) *Lowell and Mars*. University of Arizona Press, Tuscon, AZ 1231

Huygens C (1698) *Kosmotheōros, sive de terris celestibus, earumque ornatu, conjecturæ*. Den Haag 1232
1233

Irwin NI, Schulze-Makuch D (2001) Assessing the plausibility of life on other worlds. *Astrobiology* 1:143–160 1234
1235

Joyce GF (1994) Foreword. In: Deamer DW, Fleischaker GR (eds) *Origins of life: the central concepts*. Jones & Bartlett, Boston, MA, pp xi–xii 1236
1237

Kane SR, Gelino DM (2012) The habitable zone and extreme planetary orbits. *Astrobiology* 12:940–945 1238
1239

Klein HP, Horowitz NH, Levin GV, Oyama VI, Lederberg J, Rich A, Hubbard JS, Hobby GL, Straat PA, Berdahl BJ, Carle GC, Brown FS, Johnson RD (1976) The Viking biological investigation: preliminary results. *Science* 194:99–105 1240
1241
1242

Kooijman SALM (2004) On the co-evolution of life and its environment. In: Schneider SH, Miller JR, Crist E et al (eds) *Scientists debate Gaia: the next century*. MIT Press, Cambridge, MA, pp 343–351 1243
1244
1245

Kraus J (1979) We wait and wonder. *Cosmic Search* 1:31–34 1246

Lakoff G (1990) *Women, fire, and dangerous things: what categories reveal about the mind*. University of Chicago Press, Chicago, IL 1247
1248

Lederberg J (1965) Signs of life: criterion system of exobiology. *Nature* 207:9–13 1249

Levin GV, Straat PA (1976) Viking labeled release biology experiment: interim results. *Science* 194:1322–1329 1250
1251

Levin GV, Straat PA (1979) Completion of the Viking labeled release experiment on Mars. *J Mol Evol* 14:167–183 1252
1253

Levin GV, Straat PA (2016) The case for extant life on Mars and its possible detection by the Viking labeled release experiment. *Astrobiology* 16:798–810 1254
1255

Lin HW, Abad GG, Loeb A (2014) Detecting industrial pollutants in the atmospheres of Earth-like exoplanets. *Astrophys J* 792 1256
1257

Losch A (2017) *What is life?: on Earth and beyond*. Cambridge University Press, Cambridge 1258

Loveck JE (1965) A physical basis for life detection experiments. *Nature* 207:568–570 1259

Lowell P (1895) *Mars*. Houghton, Mifflin, Boston, MA 1260

Luisi P (1998) About various definitions of life. *Orig Life Evol Biosph* 28:613–622 1261

Maor E (2000) *June 8, 2004: Venus in transit*. Princeton University Press, Princeton, NJ 1262

Marov MY (2005) Mikhail Lomonosov and the discovery of the atmosphere of Venus during the 1761 transit. In: Kurtz DW (ed) *Transits of Venus: new views of the solar system and galaxy*. Cambridge University Press, Cambridge 1263
1264
1265

McKay CP (2007) How to search for life on other worlds. In: Sullivan WT, Baross JA (eds) *Planets and life: the emerging science of astrobiology*. University Press Cambridge, Cambridge, pp 461–472 1266
1267
1268

McKay DS, Gibson EK Jr, Thomas-Keptra KL et al (1996) Search for past life on Mars: possible relic biogenic activity in Martian meteorite ALH84001. *Science* 273:924–930 1269
1270

Moore P (1956) *The planet Venus*. Faber and Faber, London 1271

Oliver J, Perry RS (2006) Definitely life but not definitively. *Orig Life Evol Biosph* 36:515–521 1272

Palyi G, Zucchi C, Caglioti L (2002) *Fundamentals of life*. Elsevier, New York 1273

Peirce CS (1932) *Collected papers 2: elements of logic*. Belknap Press of Harvard University Press, Cambridge, MA 1274
1275

Peirce CS (1957) *Essays in the philosophy of science*. In: Tomas V (ed), *Liberal Arts Press*, New York 1276
1277

Peirce CS (1906/1998) *The basis of pragmatism in the normative sciences*. In: *The essential Peirce: selected philosophical writings, vol 2*. Indiana University Press, Bloomington, IN, pp 1893–1913 1278
1279
1280

- 1281 Pennock RT (2012) Negotiating boundaries in the definition of life: Wittgensteinian and Darwinian
1282 insights on resolving conceptual border conflicts. *Synthese* 185:5–20
- 1283 Persson E (2013) Philosophical aspects of astrobiology. In: Dunér D, Parthemore J, Persson E et al
1284 (eds) *The history and philosophy of astrobiology: perspectives on extraterrestrial life and the*
1285 *human mind*. Cambridge Scholars Publishing, Newcastle-upon-Tyne, pp 29–48
- 1286 Pilcher CB (2003) Biosignatures of early earths. *Astrobiology* 3:471–486
- 1287 Popper K (1963) *Conjectures and refutations: the growth of scientific knowledge*. Routledge,
1288 London
- 1289 Proctor R (1874) *Transits of Venus: a popular account of past and coming transits from the first*
1290 *observed by Horrocks A.D. 1639 to the transit of A.D. 2012*. Longmans, Green, London
- 1291 Pross A (2012) *What is life?: how chemistry becomes biology*. Oxford University Press, Oxford
- 1292 Raulin Cerceau F (2013) Pioneering concepts of planetary habitability. In: Vakoch DA
1293 (ed) *Astrobiology, history, and society: life beyond earth and the impact of discovery*. Springer,
1294 Berlin, pp 115–129
- 1295 Robus O, Haydon N, McGlynn S et al (2009) Life as a functional concept: functionalism as a robust
1296 framework for theories and definitions of multi-realized living systems. *Orig Life Evol Biosph*
1297 39:390
- 1298 Romanini V, Fernández E (2014) Peirce and biosemiotics: a guess at the riddle of life. Springer,
1299 New York
- 1300 Rosch E (1975) Cognitive representations of semantic categories. *J Exp Psychol Gen* 104:192–233
- 1301 Rosch E (1978) Principles of categorization. In: Rosch E, Lloyd BB (eds) *Cognition and catego-*
1302 *rization*. Erlbaum, Hillsdale, NJ, pp 27–48
- 1303 Rosch E (1987) Wittgenstein and categorization research in cognitive psychology. In: Chapman M,
1304 Dixon R (eds) *Meaning and the growth of understanding: Wittgenstein's significance for*
1305 *developmental psychology*. Erlbaum, Hillsdale, NJ, pp 151–166
- 1306 Ruiz-Mirazo K, Peretó J, Moreno A (2004) A universal definition of life: autonomy and open-ended
1307 evolution. *Orig Life Evol Biosph* 34:323–346
- 1308 Sagan C (1973) *Communication with extraterrestrial intelligence*. MIT Press, Cambridge, MA
- 1309 Sagan C, Thompson WR, Carlson R, Gurnett D, Hord C (1993) A search for life on Earth from the
1310 Galileo spacecraft. *Nature* 365:715–721
- 1311 Saint-Gelais R (2014) Beyond linear B: the metasemiotic challenge of communication with
1312 extraterrestrial intelligence. In: Vakoch DA (ed) *Archaeology, anthropology, and interstellar*
1313 *communication*. NASA, Washington, DC, pp 78–93
- 1314 Schopf JW (1993) Microfossils of the early Archean Apex chert: new evidence of the antiquity of
1315 life. *Science* 260:640–646
- 1316 Schröter JH (1793) *Neuere Beobachtungen der Venuskugel*. *Astronomisches Jahrbuch für 1793*
- 1317 Schröter JH (1796) *Aphroditographische Fragmente, zur genauern Kenntniss des Planeten Venus*.
1318 Helmstedt
- 1319 Schuch HP (2011) Project Ozma: the birth of observational SETI. In: Schuch HP (ed) *Searching for*
1320 *extraterrestrial intelligence: SETI past, present, and future*. Springer, Berlin, pp 13–18
- 1321 Schulze-Makuch D, Irwin LN, Fairén AG (2015) *Extraterrestrial life: what are we looking for?* In:
1322 Kolb VM (ed) *Astrobiology: an evolutionary approach*. CRC Press, Boca Raton, FL, pp
1323 399–412
- 1324 Seager S (2014) The future of spectroscopic life detection on exoplanets. *Proc Natl Acad Sci USA*
1325 111:12634–12640
- 1326 Seager S, Bains W (2015) The search for signs of life on exoplanets at the interface of chemistry and
1327 planetary science. *Sci Adv* 1:1–11
- 1328 Sellers D (2001) *The transit of Venus: the quest to find the true distance of the Sun*. MagaVelda,
1329 Leeds
- 1330 Sheehan W (1988) *Planets and perception: telescopic views and interpretations, 1609–1909*.
1331 University of Arizona Press, Tucson
- 1332 Sheehan W, Dobbins T (2001) *Epic Moon: a history of lunar exploration in the age of the telescope*.
1333 Willmann-Bell, Richmond, VA

- Sheehan W, Westfall J (2004) The transits of Venus. Prometheus Books, Amherst, NY 1334
- Shostak S (2015) Fraction of civilizations that develop a technology that releases detectable signs of their existence into space, f_c , 1961 to the present. In: Vakoch DA, Dowd MF (eds) The drake equation: estimating the prevalence of extraterrestrial life through the ages. Cambridge University Press, Cambridge, pp 27–240 1335
- Sonesson G (1994) Prolegomena to the semiotic analysis of prehistoric visual displays. *Semiotica* 100:267–332 1336
- Sonesson G (2007) From the meaning of embodiment to the embodiment of meaning. In: Zimke T, Zlatev J, Frank R (eds) *Body, language and mind. Vol. 1: embodiment*. Mouton, Berlin 1341
- Sonesson G (2009) The view from Husserl's lectern: considerations on the role of phenomenology in cognitive semiotics. *Cybern Human Knowing* 16:107–148 1342
- Sonesson G (2013) Preparations for discussing constructivism with a Martian (the second coming). In: Dunér D, Parthemore J, Persson E et al (eds) *The history and philosophy of astrobiology: perspectives on the human mind and extraterrestrial life*. Cambridge Scholars Publishing, Newcastle-upon-Tyne, pp 189–204 1343
- Sonesson G, Dunér D (2016) The cognitive semiotics of cultural evolution. In: Dunér D, Sonesson GP (eds) *Human lifeworlds: the cognitive semiotics of cultural evolution*. Peter Lang, Bern 1344
- Spranzi M (2004) Galileo and the mountains of the moon: analogical reasoning, models and metaphors in scientific discovery. *J Cogn Cult* 4:451–483 1345
- Sterken C, Aspaas PP (2013) Meeting Venus: a collection of papers presented at the Venus transit conference in Tromsø 2012. *J Astronom Data* 19:145–167 1346
- Stertzik MF, Bagnulo S, Palle E (2012) Biosignatures as revealed by spectropolarimetry of earth-shine. *Nature* 483:64–66 1347
- Strauss P (2001) Percival Lowell: the culture and science of a Boston Brahmin. Harvard University Press, Cambridge, MA 1348
- Sullivan WT, Baross JA (2007) *Planets and life: the emerging science of astrobiology*. Cambridge University Press, Cambridge 1349
- Sullivan WT, Carney D (2007) History of astrobiological ideas. In: Sullivan WT, Baross JA (eds) *Planets and life: the emerging science of astrobiology*. Cambridge University Press, Cambridge, pp 9–45 1350
- Tarter JC (2001) The search for extraterrestrial intelligence (SETI). *Annu Rev Astron Astrophys* 39:511–548 1351
- Tarter JC, Backus PR, Mancinelli RL et al (2007) A reappraisal of the habitability of planets around M dwarf stars. *Astrobiology* 7:30–65 1352
- Taylor JR (2002) *Cognitive grammar*. Oxford University Press, Oxford 1353
- Taylor JR (2003) *Linguistic categorization*. Oxford University Press, Oxford 1354
- Traphagan J (2015) *Extraterrestrial intelligence and human imagination: SETI at the intersection of science, religion, and culture*. Springer, New York 1355
- Vakoch DA (1998) Constructing messages to extraterrestrials: an exosemiotic perspective. *Acta Astronaut* 42:697–704 1356
- Vakoch DA, Dowd MF (2015) The Drake equation: estimating the prevalence of extraterrestrial life through the ages. Cambridge University Press, Cambridge 1357
- Vitruvius (1934) *On architecture: Books VI–X*, ed. Granger F. Harvard University Press, Cambridge, MA 1358
- Watson AJ (1999) Coevolution of the earth's environment and life: Goldilocks, Gaia and the anthropic principle. *Geol Soc Spec Pub* 150:75–88 1359
- Westall F, Cavalazzi B (2011) Biosignatures in rocks. In: Thiel V, Reitner J (eds) *Encyclopedia of geobiology*. Springer, Berlin, pp 189–201 1360
- Westall F, Gobbi P, Gerneke D, Mazzotti G (1998) Ultrastructure in the carbonate globules of Martian meteorite ALH84001. In: Chela-Flores J, Raulin F (eds) *Exobiology: matter, energy, and information in the origin and evolution of life in the universe*. Kluwer, Amsterdam, pp 245–250 1361

- 1386 Westall F, Foucher F, Bost N et al (2015) Biosignatures on Mars: what, where, and how?
1387 Implications for the search for Martian life. *Astrobiology* 15:998–1029
- 1388 Weston A (1988) Radio astronomy as epistemology: some philosophical reflections on the con-
1389 temporary search for extraterrestrial intelligence. *Monist* 71:88–100
- 1390 Wheeler W (2006) *The whole creature: complexity, biosemiotics and the evolution of culture.*
1391 Lawrence and Wishart, London
- 1392 Wittgenstein L (1953) *Philosophical investigations.* Blackwell, Oxford University Press, Oxford
- 1393 Woolf H (1959) *The transits of Venus: a study of eighteenth-century science.* Princeton University
1394 Press, Princeton, NJ
- 1395 Wulf A (2012) *Chasing Venus: the race to measure the heavens.* William Heinemann, London
- 1396 Zlatev J (2015) Cognitive semiotics. In: Trifonas P (ed) *International handbook of semiotics.*
1397 Springer, Berlin, pp 1043–1068

Uncorrected Proof